Unsteady Motion of a Slightly Rarefied Gas Caused by a Plate Oscillating in its Normal Direction

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Unsteady motion of a rarefied gas between two parallel plates caused when one of the plates starts a harmonic oscillation in its normal direction is investigated under the slightly rarefied condition, i.e., for small Knudsen numbers. The compressible Navier–Stokes equations are employed, and their appropriate temperature jump condition is derived systematically. The equations with the correct boundary conditions are solved numerically to give the unsteady flow field. In particular, the time-periodic solution established at later times is investigated in detail, and it is shown that the one-period average of the momentum and energy transferred from the oscillating plate to the resting one take nonzero values in contrast to the linear theory. This confirms the numerical result based on the Bhatnagar–Gross–Krook model of the Boltzmann equation for intermediate Knudsen numbers [T. Tsuji and K. Aoki, Microfluid. Nanofluid. **16**, 1033 (2014)]. It is also shown that the gas approaches the time-periodic motion exponentially fast in time.

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I. INTRODUCTION

Moving boundary problems for the Boltzmann and related kinetic equations have attracted much attention in micro fluid dynamics and in rarefied gas dynamics (e.g., [1–7]) in connection with the importance of the gas motion caused by a micromechanical oscillator in MEMS devices (e.g., [8–12]) and of the wave propagation in a rarefied gas (e.g., [13–21]). When a system containing an oscillating boundary is of micro scale or in vacuum facilities, the mean free path of the gas molecules can be comparable to the size of the system. In this case, the ordinary continuum gas dynamics is not applicable and is to be replaced by kinetic theory [22–25]. In addition, even when the system is of ordinary size in an atmospheric pressure, if the oscillator makes an oscillation with a very high frequency comparable to the collision frequency of the gas molecules, the ordinary gas dynamics should be replaced by kinetic theory.

In the present paper, we consider a fundamental problem containing an oscillating boundary. More specifically, we consider the time-dependent motion of a gas between two infinitely wide plates parallel to each other when one of the plates starts a harmonic oscillation in its normal direction.

This problem, as well as the problem without the stationary plate (i.e., the gas occupies the half space bounded by the oscillating plate), is one of the most fundamental time-dependent problems in kinetic theory and has been investigated by many authors (e.g., [13–21]). However, many of the existing works are based on the linearized setting, in which the speed of the oscillating plate is much smaller than the sonic speed, and consider the time-periodic state. In the present study, we focus our attention on the fully nonlinear setting in which the speed and amplitude of the oscillation of the plate can be large, as studied in [15, 16, 20, 21].

The present problem was investigated numerically in a recent paper [21] on the basis of the Bhatnagar–Gross–Krook (BGK) model of the Boltzmann equation [26, 27], and the transient

behavior of the gas approaching a time-periodic state was clarified. In this problem, the oscillating plate creates and propagates discontinuities in the velocity distribution function of the gas molecules continuously. In [21], a numerical method that can describe the propagation of the discontinuities proposed in [4] was used, and an accurate numerical solution was obtained for a wide range of the Knudsen number, the ratio of the mean free path of the gas molecules to the characteristic length of the system. But, since the method is computationally expensive, it is hard to obtain the long-time behavior as well as the solution for small Knudsen numbers.

For small Knudsen numbers, however, it is commonly known that the Navier–Stokes equations can describe the behavior of the gas well if appropriate slip boundary conditions are used. This may give a good alternative to the Boltzmann and its model equations that enables us to investigate the long-time behavior more easily. In fact, the slip flow theory has been established by Y. Sone in a series of papers by a systematic asymptotic analysis for the Boltzmann equation [28–33] (see also [24, 25]). For ordinary solid boundaries, the theory consists of (see [24, 25]) the linear theory for small Reynolds numbers [28, 29, 31], the weakly nonlinear theory for finite Reynolds numbers [29–31], the partially nonlinear theory for finite Reynolds numbers [32], and the fully nonlinear theory for large Reynolds numbers [33], each of which provides appropriate fluid-dynamic equations, slip boundary conditions, and the corrections in the vicinity of the boundary (Knudsen layer). However, the theory is restricted to time-independent problems, so that it cannot be applied to the present problem. The linear theory has been extended to time-dependent problems recently [34–36] (see also Sec. 3.7 in [25]). But, since the boundary is assumed not to be moving in its normal direction, it is not applicable to the present problem even if we consider the linear setting.

For this reason, we need a different framework for the present problem. In the present study, we decided to use the compressible Navier–Stokes equations as our basic equation. However, if we search the appropriate and correct slip boundary conditions for the compressible Navier–Stokes equations that can be applied immediately to the present problem with a moving plate, we notice that it is hard to find them in the literature. For this reason, following the standard procedure [24, 25, 37, 38], that is, the analysis of the Knudsen layer combined with the Chapman–Enskog expansion, we try to derive the slip boundary conditions, specialized to the present problem, in a systematic way. Then, we apply the compressible Navier–Stokes equations and the derived slip conditions, consisting of the no-slip condition for the flow velocity and a jump condition for the temperature, to the present problem and investigate the unsteady behavior of the gas and the process of approach to a time-periodic state numerically.

The paper is organized as follows. After the introduction in Sec. I, we state the problem and the assumptions in Sec. II and formulate the problem using kinetic theory in Sec. III. Then, we summarize the Chapman–Enskog solution and the compressible Navier–Stokes equations in Sec. IV and the jump boundary conditions in Sec. V. Section VI is devoted to the explanation of the numerical method for the compressible Navier–Stokes equations, and the numerical results are given in Sec. VII. Section VIII contains short concluding remarks. In addition, we summarize the basic matters concerning the collision operator of the Boltzmann equation in Appendix A and derive the slip boundary conditions in Appendix B.

II. PROBLEM AND ASSUMPTIONS

Let us consider a gas between two infinitely wide plates, kept at temperature \tilde{T}_0 and parallel to each other. One of the plates is located at $\tilde{x}_1 = \tilde{a}_w$, and the other at $\tilde{x}_1 = \tilde{d} (> \tilde{a}_w)$, where $(\tilde{x}_1, \tilde{x}_2, \tilde{x}_3)$ is the Cartesian coordinate system, and the gas is in a uniform equilibrium state at rest at density $\tilde{\rho}_0$ and temperature \tilde{T}_0 . At time $\tilde{t} = 0$, the plate at $\tilde{x}_1 = \tilde{a}_w$ starts a harmonic oscillation with amplitude \tilde{a}_w and angular frequency $\tilde{\omega}$ around $\tilde{x}_1 = 0$, that is, the location of the plate is described as $\tilde{x}_1 = \tilde{x}_w(\tilde{t})$ with $\tilde{x}_w(\tilde{t}) = \tilde{a}_w \cos \tilde{\omega} \tilde{t}$ for $\tilde{t} \ge 0$ (see Fig. 1). We investigate the unsteady motion of the gas numerically under the following assumptions:

- (i) The behavior of the gas is described by the Boltzmann equation.
- (ii) The gas molecules undergo diffuse reflection on the surfaces of the plates.
- (iii) The problem is spatially one dimensional, so that the physical quantities depend on neither \tilde{x}_2 nor \tilde{x}_3 , and the macroscopic gas flow is perpendicular to the plates.



FIG. 1: A gas between an oscillating plate and a stationary plate.

(iv) The mean free path of the gas molecules is much shorter than the characteristic length of the system. That is, the gas is slightly rarefied.

Because of assumption (iv), we will analyze the problem using the Navier–Stokes equations and the appropriate jump conditions on the plates that are consistent with the assumptions (i) and (ii).

III. FORMULATION OF THE PROBLEM USING KINETIC THEORY

A. Notations

Before formulating the problem, we introduce some notations. Let ζ_i be the velocity of gas molecules, $\tilde{f}(\tilde{t}, \tilde{x}_1, \tilde{\zeta}_i)$ the velocity distribution function of gas molecules, $\tilde{\rho}(\tilde{t}, \tilde{x}_1)$ the mass density of the gas, $\tilde{v}_i(\tilde{t}, \tilde{x}_1) = (\tilde{v}_1(\tilde{t}, \tilde{x}_1), 0, 0)$ the flow velocity of the gas [cf. assumption (iii)], $\tilde{T}(\tilde{t}, \tilde{x}_1)$ the temperature of the gas, and $\tilde{\rho}(\tilde{t}, \tilde{x}_1) = R\tilde{\rho}\tilde{T}$ the pressure of the gas, where R is the gas constant per unit mass $(R = k_B/m \text{ with the Boltzmann constant } k_B$ and the mass of a gas molecule m). In addition, let $\tilde{p}_{ij}(\tilde{t}, \tilde{x}_1)$ be the stress tensor and $\tilde{q}_i(\tilde{t}, \tilde{x}_1)$ the heat-flow vector. We set the reference time \tilde{t}_0 , the reference speed \tilde{c}_0 , and the reference length L as

$$\tilde{t}_0 = 1/\tilde{\omega}, \qquad \tilde{c}_0 = (2R\tilde{T}_0)^{1/2}, \qquad L = \tilde{c}_0 \tilde{t}_0 = (2R\tilde{T}_0)^{1/2}/\tilde{\omega}.$$
(1)

Then, we introduce the dimensionless quantities $(t, x_i, \zeta_i, f, \rho, v_i, T, p, p_{ij}, q_i)$ that correspond to $(\tilde{t}, \tilde{x}_i, \tilde{\zeta}_i, \tilde{f}, \tilde{\rho}, \tilde{v}_i, \tilde{T}, \tilde{p}, \tilde{p}_{ij}, \tilde{q}_i)$ by the following relations:

$$t = \tilde{t}/\tilde{t}_{0} = \tilde{t}\tilde{\omega}, \qquad x_{i} = \tilde{x}_{i}/L, \qquad \zeta_{i} = \tilde{\zeta}_{i}/\tilde{c}_{0}, f(t, x_{1}, \zeta_{i}) = (\tilde{c}_{0}^{3}/\tilde{\rho}_{0})\tilde{f}(\tilde{t}, \tilde{x}_{1}, \tilde{\zeta}_{i}), \rho(t, x_{1}) = \tilde{\rho}/\tilde{\rho}_{0}, \qquad v_{i}(t, x_{1}) = (v_{1}(t, x_{1}), 0, 0) = \tilde{v}_{i}/\tilde{c}_{0},$$
(2)
$$T(t, x_{1}) = \tilde{T}/\tilde{T}_{0}, \qquad p(t, x_{1}) = \tilde{p}/\tilde{p}_{0}, p_{ij}(t, x_{1}) = \tilde{p}_{ij}/\tilde{p}_{0}, \qquad q_{i}(t, x_{1}) = \tilde{q}_{i}/\tilde{\rho}_{0}\tilde{c}_{0},$$

where $\tilde{p}_0 = R\tilde{\rho}_0\tilde{T}_0$ is the reference pressure based on $\tilde{\rho}_0$ and \tilde{T}_0 . Then, the dimensionless macroscopic quantities ρ , v_i , T, p, p_{ij} , and q_i are expressed as follows [25]:

$$\rho = \int f d\boldsymbol{\zeta}, \qquad v_i = \frac{1}{\rho} \int \zeta_i f d\boldsymbol{\zeta}, \qquad T = \frac{2}{3\rho} \int (\zeta_i - v_i)^2 f d\boldsymbol{\zeta}, \qquad p = \rho T, \qquad (3a)$$

$$p_{ij} = 2 \int (\zeta_i - v_i)(\zeta_j - v_j) f d\boldsymbol{\zeta}, \qquad q_i = \int (\zeta_i - v_i)(\zeta_j - v_j)^2 f d\boldsymbol{\zeta}, \tag{3b}$$

where $d\zeta = d\zeta_1 d\zeta_2 d\zeta_3$. Here and in what follows, the summation convention (i.e., $a_i b_i = a_1 b_1 + a_2 b_2 + a_3 b_3$, $a_i^2 = a_1^2 + a_2^2 + a_3^2$) is used, and the domain of integration with respect to ζ_i is its whole space unless otherwise stated. The assumption $v_2 = v_3 = 0$ [assumption (iii)] means

that we implicitly assume that f is cylindrically symmetric in ζ_i , i.e., $f = f(t, x_1, \zeta_1, \sqrt{\zeta_2^2 + \zeta_3^2})$. Therefore, some components of p_{ij} and q_i vanish, that is,

$$p_{12} = p_{21} = p_{23} = p_{32} = p_{31} = p_{13} = 0, \qquad q_2 = q_3 = 0.$$
(4)

Let us denote by \tilde{v}_w the velocity (in the \tilde{x}_1 direction) of the oscillating plate, i.e., $\tilde{v}_w = -\tilde{a}_w \tilde{\omega} \sin \tilde{\omega} \tilde{t}$. We introduce the dimensionless position $x_w(t)$ and velocity $v_w(t)$ of the oscillating plate and the dimensionless parameters a_w and d corresponding to the amplitude \tilde{a}_w of the oscillating plate and the position \tilde{d} of the resting plate by

$$x_w(t) = \tilde{x}_w/L, \qquad v_w(t) = \tilde{v}_w/\tilde{c}_0, \qquad a_w = \tilde{a}_w/L, \qquad d = d/L.$$
 (5)

Then

$$x_w(t) = a_w \cos t, \qquad v_w(t) = -a_w \sin t. \tag{6}$$

It should be noted that the dimensionless amplitude (and speed) of the plate a_w is of the order of the Mach number Ma based on the maximum speed of the plate, that is,

$$a_w = \sqrt{5/6} \,\mathrm{Ma}, \qquad \mathrm{Ma} = \tilde{a}_w \tilde{\omega} / [(5/3) R \widetilde{T}_0]^{1/2}.$$
 (7)

B. Basic equations

The dimensionless Boltzmann equation [25] in the present spatially one-dimensional problem is written as

$$\frac{\partial f}{\partial t} + \zeta_1 \frac{\partial f}{\partial x_1} = \frac{1}{\epsilon} J(f, f), \tag{8}$$

with a small parameter ϵ :

$$\epsilon = (\sqrt{\pi}/2) \operatorname{Kn} = (\sqrt{\pi}/2) (l_0/L), \qquad (9)$$

where Kn is the Knudsen number, and l_0 is the mean free path of the gas molecules at the initial equilibrium state at rest. In Eq. (8), J(f, f) is the dimensionless collision term the explicit form of which is given in Appendix A.

The diffuse-reflection condition [25] on the moving boundary is written as

$$f(t, x_w(t), \zeta_i) = \frac{\sigma_w}{\pi^{3/2}} \exp\left(-\left\{ [\zeta_1 - v_w(t)]^2 + \zeta_2^2 + \zeta_3^2 \right\} \right), \quad \text{for } \zeta_1 - v_w(t) > 0, \tag{10a}$$

$$\sigma_w = -2\pi^{1/2} \int_{\zeta_1 - v_w(t) < 0} [\zeta_1 - v_w(t)] f(t, x_w(t), \zeta_i) d\boldsymbol{\zeta},$$
(10b)

where x_w and v_w are given by Eq. (6), and that on the plate at rest is

$$f(t, d, \zeta_i) = \frac{\sigma_w}{\pi^{3/2}} \exp\left(-(\zeta_1^2 + \zeta_2^2 + \zeta_3^2)\right), \quad \text{for } \zeta_1 < 0, \tag{11a}$$

$$\sigma_w = 2\pi^{1/2} \int_{\zeta_1 > 0} \zeta_1 f(t, \, d, \, \zeta_i) d\boldsymbol{\zeta}.$$
 (11b)

The initial condition is

$$f(0, x_1, \zeta_i) = \pi^{-3/2} \exp\left(-(\zeta_1^2 + \zeta_2^2 + \zeta_3^2)\right), \quad \text{for } a_w \le x_1 \le d.$$
(12)

It is seen from Eqs. (8), (10a), (11a), and (12) that the problem is characterized by the three parameters ϵ , a_w , and d.

IV. CHAPMAN-ENSKOG SOLUTION AND THE COMPRESSIBLE NAVIER-STOKES EQUATIONS

The Chapman–Enskog expansion [25, 39, 40] is a well-known procedure to derive the Euler and Navier–Stokes equations for a compressible fluid from the Boltzmann equation. In this section, we summarize the Chapman–Enskog solution obtained by the Chapman–Enskog expansion and the resulting compressible Navier–Stokes equations for the present spatially one-dimensional problem. We basically follow the description and notation of Sec. B4 in [25], but note that some symbols are not exactly the same. For example, the hat $\hat{}$ indicating the dimensionless quantities in [25] is omitted here, and C_i in [25] is denoted by C_i here.

The first-order Chapman–Enskog solution for the spatially one-dimensional Boltzmann equation (8) can be expressed as

$$f = f_{\rm CE}^{(1)} + O(\epsilon^2) = f^{(0)} + f^{(1)}\epsilon + O(\epsilon^2).$$
(13)

Here, the leading-order term $f^{(0)}$ is a local Maxwellian distribution

$$f^{(0)} = \frac{\rho}{(\pi T)^{3/2}} \exp\left(-\frac{(\zeta_1 - v_1)^2 + \zeta_2^2 + \zeta_3^2}{T}\right) = \frac{\rho}{T^{3/2}} E(C), \tag{14}$$

and the first-order term $f^{(1)}$ takes the following form:

$$f^{(1)} = f^{(0)}\Psi,$$
 (15a)

$$\Psi = -\frac{1}{\rho T^{1/2}} \left(C_1^2 - \frac{1}{3} C^2 \right) \frac{\partial v_1}{\partial x_1} \mathcal{B}^{(0)}(C, T) - \frac{1}{\rho T} C_1 \frac{\partial T}{\partial x_1} \mathcal{A}(C, T),$$
(15b)

where

$$C_1 = \frac{\zeta_1 - v_1}{T^{1/2}}, \qquad C_2 = \frac{\zeta_2}{T^{1/2}}, \qquad C_3 = \frac{\zeta_3}{T^{1/2}},$$
 (16a)

$$E(C) = \pi^{-3/2} \exp(-C^2), \qquad C = (C_1^2 + C_2^2 + C_3^2)^{1/2},$$
 (16b)

and $\mathcal{B}^{(0)}(C, T)$ and $\mathcal{A}(C, T)$ are, in principle, known functions of C and T, which are specified in Appendix A. The expansion (13) is designed in such a way that the first-order term $f^{(1)}$ satisfies the constraint:

$$\int (1, \,\zeta_1, \,\zeta_i^2) f^{(1)} d\boldsymbol{\zeta} = 0, \tag{17}$$

and the same is true for the $O(\epsilon^2)$ term in Eq. (13). Therefore, ρ , v_1 , and T contained in $f^{(0)}$ are nothing but the density, the flow velocity in the x_1 direction, and the temperature associated with f of the expansion (13) [cf. Eq. (3a)].

With Eq. (13), p_{ij} and q_i in Eq. (3b) become

$$p_{11} = p - \frac{4}{3}\epsilon \Gamma_1(T)\frac{\partial v_1}{\partial x_1}, \qquad p_{22} = p_{33} = p, \qquad p_{ij} = 0 \quad (i \neq j), \tag{18a}$$

$$q_1 = -\frac{5}{4}\epsilon \,\Gamma_2(T)\frac{\partial T}{\partial x_1}, \qquad q_2 = q_3 = 0, \tag{18b}$$

where, $\Gamma_1(T)$ and $\Gamma_2(T)$ are expressed as

$$\Gamma_1(T) = \frac{8}{15\sqrt{\pi}} T^{1/2} \int_0^\infty C^6 \mathcal{B}^{(0)}(C, T) e^{-C^2} dC, \qquad (19a)$$

$$\Gamma_2(T) = \frac{16}{15\sqrt{\pi}} T^{1/2} \int_0^\infty C^6 \mathcal{A}(C, T) e^{-C^2} dC,$$
(19b)

and are related to the viscosity μ and the thermal conductivity λ as

$$\mu = (\tilde{p}_0 L/\tilde{c}_0) \,\epsilon \,\Gamma_1(T) = (\sqrt{\pi}/2)(\tilde{p}_0 l_0/\tilde{c}_0) \,\Gamma_1(T), \tag{20a}$$

$$\lambda = (5/4)(\tilde{p}_0 \tilde{c}_0 L/\tilde{T}_0) \epsilon \Gamma_2(T) = (5\sqrt{\pi}/8)(\tilde{p}_0 \tilde{c}_0 l_0/\tilde{T}_0) \Gamma_2(T).$$
(20b)

For hard-sphere molecules and the BGK model, $\Gamma_1(T)$ and $\Gamma_2(T)$ are explicitly expressed as

$$\Gamma_1(T) = 1.270042427 T^{1/2}, \quad \Gamma_2(T) = 1.922284066 T^{1/2} \quad \text{(hard sphere)}, \tag{21a}$$

$$\Gamma_1(T) = \Gamma_2(T) = T \quad \text{(BGK)}. \tag{21b}$$

With Eq. (18), the Maxwell transport equations, which are derived by integrating Eq. (8) times $(1, \zeta_i, \zeta_i^2)$ over the whole space of ζ_i , reduce to the compressible Navier–Stokes equations:

$$\frac{\partial \rho}{\partial t} + \frac{\partial (\rho v_1)}{\partial x_1} = 0, \tag{22a}$$

$$\frac{\partial(\rho v_1)}{\partial t} + \frac{\partial(\rho v_1^2)}{\partial x_1} = -\frac{1}{2}\frac{\partial p}{\partial x_1} + \frac{2}{3}\epsilon\frac{\partial}{\partial x_1}\left[\Gamma_1(T)\frac{\partial v_1}{\partial x_1}\right],\tag{22b}$$

$$\frac{\partial}{\partial t} \left[\rho \left(\frac{3}{2} T + v_1^2 \right) \right] + \frac{\partial}{\partial x_1} \left[\rho v_1 \left(\frac{5}{2} T + v_1^2 \right) \right] \\ = \frac{5}{4} \epsilon \frac{\partial}{\partial x_1} \left[\Gamma_2(T) \frac{\partial T}{\partial x_1} \right] + \frac{4}{3} \epsilon \frac{\partial}{\partial x_1} \left[\Gamma_1(T) \frac{\partial v_1}{\partial x_1} v_1 \right],$$
(22c)

where $p = \rho T$ [Eq. (3a)], and the x_2 and x_3 components of the momentum equation become trivial (i.e., 0 = 0).

If we substitute Eq. (13) with ρ , v_1 , and T satisfying the Navier–Stokes equations (22) into the Boltzmann equation (8) multiplied by ϵ , then we observe that the expansion (13) satisfies the latter with the error of $O(\epsilon^2)$.

V. JUMP BOUNDARY CONDITIONS

In the first-order Chapman–Enskog solution (13), which corresponds to the Navier–Stokes equations (22), the boundary conditions on the plates (10a) and (11a) are not taken into account. To be consistent with the fact that the term up to $O(\epsilon)$ is considered in Eq. (13), we need to satisfy the boundary conditions up to the order of ϵ .

We first try to satisfy the boundary conditions with the Chapman–Enskog solution (13). If we recall that the leading-order term $f^{(0)}$ is a local Maxwellian [Eq. (14)], it can be made to satisfy Eqs. (10a) and (11a) by assuming that

$$v_1 = v_w(t), \qquad T = 1, \qquad \text{at } x_1 = x_w(t),$$
 (23a)

$$v_1 = 0, \qquad T = 1, \qquad \text{at } x_1 = d.$$
 (23b)

On the other hand, in order that the first-order term $f^{(1)}$ satisfies Eqs. (10a) and (11a) at the order of ϵ , we need to impose the following conditions:

$$\frac{\partial v_1}{\partial x_1} = 0, \qquad \frac{\partial T}{\partial x_1} = 0, \qquad \text{at } x_1 = x_w(t) \text{ and } d.$$
 (24)

However, the constraints on the plates, Eqs. (23) and (24), are too many as the boundary conditions for the Navier–Stokes equations (22). Therefore, this approach does not work. But, the fact that the choice (23) works at the zeroth order in ϵ suggests that

$$v_1 - v_w(t) = O(\epsilon), \qquad T - 1 = O(\epsilon), \qquad \text{at } x_1 = x_w(t),$$
 (25a)

$$v_1 = O(\epsilon), \qquad T - 1 = O(\epsilon), \qquad \text{at } x_1 = d.$$
 (25b)

In order to obtain the solution satisfying the boundary conditions, it is required to introduce the kinetic boundary layer, the so-called Knudsen layer, with thickness of the order of ϵ adjacent to the plates. To be more specific, we seek the solution in the following form:

$$f = f_{\rm CE}^{(1)} + f^{(0)}\Phi\epsilon + O(\epsilon^2) = f^{(0)}(1 + \Psi\epsilon + \Phi\epsilon) + O(\epsilon^2),$$
(26)

where we use Eqs. (14) and (15), and Φ is the correction inside the Knudsen layer. We note that Φ has a length scale of variation of the order of ϵ in the direction normal to the boundary and vanishes outside the Knudsen layer. The correction is introduced only at the order of ϵ because, as we saw, the leading-order term $f^{(0)}$ of the Chapman–Enskog expansion could be made to satisfy the boundary conditions.

The analysis of the Knudsen-layer correction Φ determines the boundary conditions for the Navier–Stokes equations (22) that are correct up to the order of ϵ . Leaving the detailed analysis of Φ and the process of deriving the boundary conditions in Appendix B, we summarize the resulting boundary conditions here, that is,

$$v_1 - v_w(t) = 0, \qquad T - 1 = \frac{1}{\rho} \frac{\partial v_1}{\partial x_1} \alpha_v \epsilon + \frac{1}{\rho} \frac{\partial T}{\partial x_1} \alpha_T \epsilon \qquad \text{at} \quad x_1 = x_w(t),$$
 (27a)

$$v_1 = 0, \qquad T - 1 = \frac{1}{\rho} \frac{\partial v_1}{\partial x_1} \alpha_v \epsilon - \frac{1}{\rho} \frac{\partial T}{\partial x_1} \alpha_T \epsilon \qquad \text{at } x_1 = d,$$
 (27b)

where α_v and α_T are constants that depend on the model for the intermolecular collision. For hard-sphere molecules and for the BGK model, they are given as

$$\alpha_v = 0.45957, \qquad \alpha_T = 2.4001, \qquad \text{(hard sphere)}, \tag{28a}$$

$$\alpha_v = 0.44045, \qquad \alpha_T = 1.30272, \qquad (BGK).$$
 (28b)

These values are for the diffuse reflection condition. For different boundary conditions of ordinary type, these constants take different values, the form of the boundary conditions (27) being unchanged. Although the velocity v_1 is the same as that of the plate, the temperature T differs from that of the plates 1, that is, there is a *temperature jump* of $O(\epsilon)$. The term containing α_T is the usual temperature jump proportional to the temperature gradient normal to the plates. The term containing α_v is the jump proportional to the normal stress, which becomes higher order in the steady problems with small Mach number flows. We note that [15] contains the numerical analysis of the present problem using the compressible Navier–Stokes equations with slip boundary conditions. However, the temperature jump condition used there includes neither the term proportional to the normal stress nor the factor $1/\rho$ [cf. Eq. (27)]. In the linearized problem, ρ is replaced by its reference value $\rho = 1$, so that the factor $1/\rho$ does not appear. But in the nonlinear problem, there is a significant density change in the gas, and the temperature jump should be proportional to the local Knudsen number, not the global one. Since the local Knudsen number is expressed as ϵ/ρ , the factor $1/\rho$ appears in Eq. (27).

The initial condition for the Navier–Stokes equations corresponding to that for the Boltzmann equation, Eq. (12), is

$$\rho = 1, \quad v_1 = 0, \quad T = 1, \quad \text{at } t = 0.$$
 (29)

We are going to solve the Navier–Stokes equations (22) numerically under the boundary conditions (27) and the initial condition (29). In the present problem, the initial condition (12) for the Boltzmann equation is consistent with the first-order Chapman–Enskog solution (13) and satisfies the boundary conditions (10a) and (11a) at t = 0. In addition, the left plate starts the motion gradually with the fluid-dynamic time scale. Therefore, the initial layer does not appear. In other words, the Navier–Stokes equations with the initial condition (29) describe the initial stage of the gas motion correctly.

VI. NUMERICAL METHOD

In this section, we explain the numerical method used in the computation of Eqs. (22), (27), and (29). For brevity, we omit the subscript 1 of x_1 and v_1 in the present section, i.e., we let

$$x = x_1, \qquad v = v_1. \tag{30}$$

A. Numerical scheme

Let us divide the region of the gas $x_w(t) \le x \le d$ into N cells by the grid points $x_j(t)$ (j = 0, ..., N) with $x_0(t) = x_w(t)$ and $x_N(t) = d$, which move with time. In the case of uniform cells, $x_j(t) = x_w(t) + (j/N)[d - x_w(t)]$. We call the interval $x_{j-1}(t) \le x \le x_j(t)$ the *j*th cell and define the center of the *j*th cell $x_{j-\frac{1}{2}}(t)$ as the middle point, i.e., $x_{j-\frac{1}{2}}(t) = [x_{j-1}(t) + x_j(t)]/2$. We consider the spatially one-dimensional equation of the following form:

$$\frac{\partial F}{\partial t} + \frac{\partial G}{\partial x} = 0, \tag{31}$$

where, F(t, x) and G(t, x) are functions of x and t. For the Navier–Stokes equations, F and G are given by

$$F = \rho, \qquad G = \rho v, \tag{32}$$

for Eq. (22a),

$$F = \rho v, \qquad G = \rho v^2 + \frac{1}{2}p - \frac{2}{3}\epsilon \Gamma_1(T)\frac{\partial v}{\partial x}, \qquad (33)$$

for Eq. (22b), and

$$F = \rho v^2 + \frac{3}{2}\rho T, \qquad G = \left(\rho v^2 + \frac{5}{2}\rho T\right)v - \frac{5}{4}\epsilon\Gamma_2(T)\frac{\partial T}{\partial x} - \frac{4}{3}\epsilon\Gamma_1(T)v\frac{\partial v}{\partial x}, \tag{34}$$

for Eq. (22c). We try to discretize this equation on the system of moving grids $x_i(t)$.

Let us integrate Eq. (31) over the interval $[x_{j-1}(t), x_j(t)]$ in x and then $[t, t + \Delta t]$ in t:

$$\int_{t}^{t+\Delta t} \int_{x_{j-1}(\tau)}^{x_{j}(\tau)} \frac{\partial F}{\partial \tau} dx d\tau + \int_{t}^{t+\Delta t} \int_{x_{j-1}(\tau)}^{x_{j}(\tau)} \frac{\partial G}{\partial x} dx d\tau = 0.$$
(35)

The second term on the left-hand side is transformed as follows:

$$\int_{t}^{t+\Delta t} \int_{x_{j-1}(\tau)}^{x_{j}(\tau)} \frac{\partial G}{\partial x} dx d\tau = \int_{t}^{t+\Delta t} [G(\tau, x_{j}(\tau)) - G(\tau, x_{j-1}(\tau))] d\tau$$

$$\simeq [G(t, x_{j}(t)) - G(t, x_{j-1}(t))] \Delta t$$

$$\simeq [G(t + \Delta t, x_{j}(t + \Delta t)) - G(t + \Delta t, x_{j-1}(t + \Delta t))] \Delta t$$

$$\simeq [G(t, x_{j}(t + \Delta t)) - G(t, x_{j-1}(t + \Delta t))] \Delta t.$$
(36)

On the other hand, we have the following identity:

$$\frac{\partial}{\partial \tau} \int_{x_{j-1}(\tau)}^{x_j(\tau)} F(\tau, x) dx = \int_{x_{j-1}(\tau)}^{x_j(\tau)} \frac{\partial F}{\partial \tau}(\tau, x) dx + F(\tau, x_j(\tau)) \dot{x}_j(\tau) - F(\tau, x_{j-1}(\tau)) \dot{x}_{j-1}(\tau),$$
(37)

where $\dot{x}_j(t) = dx_j(t)/dt$. Using this identity, we can transform the first term on the left-hand side of Eq. (35) as

$$\int_{t}^{t+\Delta t} \int_{x_{j-1}(\tau)}^{x_{j}(\tau)} \frac{\partial F}{\partial \tau} dx d\tau$$

$$= \int_{t}^{t+\Delta t} \frac{\partial}{\partial \tau} \int_{x_{j-1}(\tau)}^{x_{j}(\tau)} F(\tau, x) dx d\tau$$

$$- \int_{t}^{t+\Delta t} [F(\tau, x_{j}(\tau))\dot{x}_{j}(\tau) - F(\tau, x_{j-1}(\tau))\dot{x}_{j-1}(\tau)] d\tau$$

$$= \int_{x_{j-1}(t+\Delta t)}^{x_{j}(t+\Delta t)} F(t+\Delta t, x) dx - \int_{x_{j-1}(t)}^{x_{j}(t)} F(t, x) dx$$

$$- \int_{t}^{t+\Delta t} [F(\tau, x_{j}(\tau))\dot{x}_{j}(\tau) - F(\tau, x_{j-1}(\tau))\dot{x}_{j-1}(\tau)] d\tau.$$
(38)

With the middle point $x_{j-\frac{1}{2}}(t)$ between $x_{j-1}(t)$ and $x_j(t)$, the first two integrals on the last line of Eq. (38) are approximated as

$$\int_{x_{j-1}(t+\Delta t)}^{x_{j}(t+\Delta t)} F(t+\Delta t, x) dx - \int_{x_{j-1}(t)}^{x_{j}(t)} F(t, x) dx$$

$$\simeq F(t+\Delta t, x_{j-\frac{1}{2}}(t+\Delta t)) [x_{j}(t+\Delta t) - x_{j-1}(t+\Delta t)]$$

$$- F(t, x_{j-\frac{1}{2}}(t)) [x_{j}(t) - x_{j-1}(t)].$$
(39)

The last integral on the last line of Eq. (38) is approximated as

$$\int_{t}^{t+\Delta t} [F(\tau, x_{j}(\tau))\dot{x}_{j}(\tau) - F(\tau, x_{j-1}(\tau))\dot{x}_{j-1}(\tau)]d\tau$$

$$\simeq F(t, x_{j}(t))[x_{j}(t+\Delta t) - x_{j}(t)] - F(t, x_{j-1}(t))[x_{j-1}(t+\Delta t) - x_{j-1}(t)].$$
(40)

Let us denote by t_k the discretized time variable. If the time step is uniform, $t_{k+1} = t_k + \Delta t$ for any k. We use the following notation:

$$x_i^k = x_i(t_k), \qquad x_w^k = x_w(t_k), \qquad F_i^k = F(t_k, x_i^k), \qquad G_i^k = G(t_k, x_i^k).$$
 (41)

If we let $t = t_k$ and $t + \Delta t = t_{k+1}$ in Eq. (35) with the approximations (36) (the approximation in the second line), (38), (39), and (40), we have the following discretized version of (31):

$$F_{j-\frac{1}{2}}^{k+1}(x_j^{k+1} - x_{j-1}^{k+1}) - F_{j-\frac{1}{2}}^k(x_j^k - x_{j-1}^k) - F_j^k(x_j^{k+1} - x_j^k) + F_{j-1}^k(x_{j-1}^{k+1} - x_{j-1}^k) + (G_j^k - G_{j-1}^k)\Delta t = 0.$$
(42)

From this equation, $F_{j-\frac{1}{2}}^{k+1}$ is expressed as

$$F_{j-\frac{1}{2}}^{k+1} = \frac{F_{j-\frac{1}{2}}^{k}(x_{j}^{k} - x_{j-1}^{k}) + F_{j}^{k}(x_{j}^{k+1} - x_{j}^{k}) - F_{j-1}^{k}(x_{j-1}^{k+1} - x_{j-1}^{k})}{x_{j}^{k+1} - x_{j-1}^{k+1}} - \frac{G_{j}^{k} - G_{j-1}^{k}}{x_{j}^{k+1} - x_{j-1}^{k+1}}\Delta t,$$
(43)

in terms of F and G at the previous time step t_k . If we use the approximation in the last line of Eq. (36) instead of that in the second line, we have

$$F_{j-\frac{1}{2}}^{k+1} = \frac{F_{j-\frac{1}{2}}^{k}(x_{j}^{k} - x_{j-1}^{k}) + F_{j}^{k}(x_{j}^{k+1} - x_{j}^{k}) - F_{j-1}^{k}(x_{j-1}^{k+1} - x_{j-1}^{k})}{x_{j}^{k+1} - x_{j-1}^{k+1}} - \frac{G(t_{k}, x_{j}^{k+1}) - G(t_{k}, x_{j-1}^{k+1})}{x_{j}^{k+1} - x_{j-1}^{k+1}}\Delta t.$$

$$(44)$$

Before moving on to the details for the Navier–Stokes equations, we interpret the meaning of Eq. (44). Let us consider the integral $\int_{x_{j-1}^{k+1}}^{x_{j}^{k+1}} F(t_k, x) dx$ and write it as

$$\int_{x_{j-1}^{k+1}}^{x_{j}^{k+1}} F(t_k, x) dx = \int_{x_{j-1}^{k}}^{x_j^k} F(t_k, x) dx + \int_{x_j^k}^{x_j^{k+1}} F(t_k, x) dx - \int_{x_{j-1}^{k}}^{x_{j-1}^{k+1}} F(t_k, x) dx.$$
(45)

We approximate this equality using the middle points as

$$F(t_k, x_{j-\frac{1}{2}}^{k+1})(x_j^{k+1} - x_{j-1}^{k+1}) \simeq F(t_k, x_{j-\frac{1}{2}}^k)(x_j^k - x_{j-1}^k) + F(t_k, x_j^k)(x_j^{k+1} - x_j^k) - F(t_k, x_{j-1}^k)(x_{j-1}^{k+1} - x_{j-1}^k).$$

$$(46)$$

Then we have,

$$F(t_k, x_{j-\frac{1}{2}}^{k+1}) \simeq \frac{F_{j-\frac{1}{2}}^k(x_j^k - x_{j-1}^k) + F_j^k(x_j^{k+1} - x_j^k) - F_{j-1}^k(x_{j-1}^{k+1} - x_{j-1}^k)}{x_j^{k+1} - x_{j-1}^{k+1}}.$$
 (47)

This expression is the same as the first fractional term on the right-hand side of Eq. (43) or (44). This gives an interpolation, using F_{j-1}^k , $F_{j-\frac{1}{2}}^k$, and F_j^k , to give the value of F at time t_k and at position $x_{j-\frac{1}{2}}^{k+1}$, i.e., at the middle point that is supposed to be at time t_{k+1} . In other words, it is an interpolation based on the conservation of F at $t = t_k$ when the *j*th cell moves from $[x_{j-1}^k, x_j^k]$ to $[x_{j-1}^{k+1}, x_j^{k+1}]$ [cf. Eq. (45)]. For later convenience, we rewrite Eqs. (44) and (47) in the following form:

$$F_{j-\frac{1}{2}}^{k+1} = F(t_k, x_{j-\frac{1}{2}}^{k+1}) - \frac{G(t_k, x_j^{k+1}) - G(t_k, x_{j-1}^{k+1})}{x_j^{k+1} - x_{j-1}^{k+1}} \Delta t,$$
(48a)

$$F(t_k, x_{j-\frac{1}{2}}^{k+1}) = \frac{F_{j-\frac{1}{2}}^k(x_j^k - x_{j-1}^k) + F_j^k(x_j^{k+1} - x_j^k) - F_{j-1}^k(x_{j-1}^{k+1} - x_{j-1}^k)}{x_j^{k+1} - x_{j-1}^{k+1}}.$$
 (48b)

Then, we can see that Eq. (48a) is a finite-difference version of Eq. (31). That is, if we consider Eq. (31) at point $x = x_{j-\frac{1}{2}}^{k+1}$ and time $t = t_k$, and if we replace $\partial F/\partial t$ with the forward difference $[F(t_{k+1}, x_{j-\frac{1}{2}}^{k+1}) - F(t_k, x_{j-\frac{1}{2}}^{k+1})]/\Delta t$ and $\partial G/\partial x$ with the central difference $[G(t_k, x_j^{k+1}) - G(t_k, x_{j-1}^{k+1})]/(x_j^{k+1} - x_{j-1}^{k+1})$, then we obtain Eq. (48a). Therefore, the scheme (48) consists of two steps: The interpolation required by grid displacement and the time marching. This method was proposed in [41, 42].

В. **Compressible Navier–Stokes equations**

Now we are ready to apply the scheme (48) to the Navier–Stokes equations, for which F and G are listed in Eqs. (32)–(34). In this subsection, we let h represent the macroscopic quantities, i.e., $h = \rho$, v, and T. Suppose that at time $t = t_k$, the quantities h_{j-1}^k , $h_{j-\frac{1}{2}}^k$ (j = 1, ..., N),

and h_N^k are all known. Then, the procedure is as follows: (i) From Eq. (48b), we obtain $F(t_k, x_{j-\frac{1}{2}}^{k+1})$, which gives $h(t_k, x_{j-\frac{1}{2}}^{k+1})$. (ii) For $1 \le j \le N-1$, we obtain the values of the macroscopic quantities at the interfaces of the new cell, $h(t_k, x_j^{k+1})$, by the linear interpolation using $h(t_k, x_{j-\frac{1}{2}}^{k+1})$ and $h(t_k, x_{j+\frac{1}{2}}^{k+1})$. For j = 0 and N (i.e., on the boundary, $x = x_0^{k+1}$ and x_N^{k+1}), we first obtain $\rho(t_k, x_j^{k+1})$ by the linear extrapolation [i.e., $\rho(t_k, x_0^{k+1}) = 2\rho(t_k, x_1^{k+1}) - \rho(t_k, x_1^{k+1})$ and $\rho(t_k, x_N^{k+1}) = 2\rho(t_k, x_1^{k+1}) - \rho(t_k, x_1^{k+1})$ $2\rho(t_k, x_{N-\frac{1}{2}}^{k+1}) - \rho(t_k, x_{N-1}^{k+1})$ and then assume that $v(t_k, x_j^{k+1})$ and $T(t_k, x_j^{k+1})$ are the same as $v(t_k, x_j^k)$ and $T(t_k, x_j^k)$, respectively [i.e., $v(t_k, x_0^{k+1}) = v(t_k, x_0^k)$, $v(t_k, x_N^{k+1}) = v(t_k, x_N^k)$, $T(t_k, x_0^{k+1}) = T(t_k, x_0^k)$, and $T(t_k, x_N^{k+1}) = T(t_k, x_N^k)$]. (iii) We obtain the pressure at the grid points $p(t_k, x_j^{k+1})$ and that at the cell centers

 $p(t_k, x_{j-\frac{1}{2}}^{k+1})$ by the equation of state $p = \rho T$.

(iv) We update the macroscopic quantities at the cell centers by the discrete equation (48a): we obtain $F_{j-\frac{1}{2}}^{k+1}$ and thus $h_{j-\frac{1}{2}}^{k+1}$. In this process, we use the finite difference based on the two points $x_{j-\frac{1}{2}}^{k+1}$ and $x_{j+\frac{1}{2}}^{k+1}$ (the central difference when the size of the cells is uniform) for $\partial v/\partial x$ and $\partial T/\partial x$ in G. For example, for G in Eq. (34), we use the following $G(t_k, x_j^{k+1})$:

$$G(t_k, x_j^{k+1}) = \left\{ \rho(t_k, x_j^{k+1}) [v(t_k, x_j^{k+1})]^2 + \frac{5}{2} \rho(t_k, x_j^{k+1}) T(t_k, x_j^{k+1}) \right\} v(t_k, x_j^{k+1}) - \frac{5}{4} \epsilon \Gamma_2(T(t_k, x_j^{k+1})) \frac{T(t_k, x_{j+\frac{1}{2}}^{k+1}) - T(t_k, x_{j-\frac{1}{2}}^{k+1})}{x_{j+\frac{1}{2}}^{k+1} - x_{j-\frac{1}{2}}^{k+1}} - \frac{4}{3} \epsilon \Gamma_1(T(t_k, x_j^{k+1})) v(t_k, x_j^{k+1}) \frac{v(t_k, x_{j+\frac{1}{2}}^{k+1}) - v(t_k, x_{j-\frac{1}{2}}^{k+1})}{x_{j+\frac{1}{2}}^{k+1} - x_{j-\frac{1}{2}}^{k+1}}.$$
(49)

(v) For $1 \leq j \leq N-1$, we obtain the interface values h_j^{k+1} by the linear interpolation using $h_{j-\frac{1}{2}}^{k+1}$ and $h_{j+\frac{1}{2}}^{k+1}$. For j = 0 and N (i.e., on the boundary, $x = x_0^{k+1}$ and x_N^{k+1}), we first obtain ρ_j^{k+1} by the linear extrapolation (i.e., $\rho_0^{k+1} = 2\rho_{\frac{1}{2}}^{k+1} - \rho_1^{k+1}$ and $\rho_N^{k+1} = 2\rho_{N-\frac{1}{2}}^{k+1} - \rho_{N-1}^{k+1}$) and then obtain v_j^{k+1} and T_j^{k+1} using the boundary conditions. More specifically, the boundary conditions, Eq. (27) discretized by using one-sided finite difference for derivatives, give the following v_j^{k+1} and T_j^{k+1} :

$$v_0^{k+1} = v_w^{k+1},\tag{50a}$$

$$\frac{1}{2}\rho_0^{k+1}(x_1^{k+1} - x_0^{k+1}) + \alpha_T \epsilon$$

$$v_N^{k+1} = 0,$$
 (50c)

$$T_N^{k+1} = \frac{\frac{1}{2}\rho_N^{k+1}(x_N^{k+1} - x_{N-1}^{k+1}) - \alpha_v \epsilon v_{N-\frac{1}{2}}^{k+1} + \alpha_T \epsilon T_{N-\frac{1}{2}}^{k+1}}{\frac{1}{2}\rho_N^{k+1}(x_N^{k+1} - x_{N-1}^{k+1}) + \alpha_T \epsilon}$$
(50d)

(vi) We obtain the pressure p_j^{k+1} and $p_{j-\frac{1}{2}}^{k+1}$ by the equation of state $p = \rho T$.

VII. RESULTS OF COMPUTATION

In this section, we show the results of numerical computation. We consider the cases of hardsphere molecules [Eqs. (21a) and (28a)] and the BGK model [Eqs. (21b) and (28b)]. We recall that the problem is characterized by three dimensionless parameters: the (modified) Knudsen number ϵ [Eq. (9)], the dimensionless amplitude of the plate a_w [Eq. (7)], and the dimensionless distance between the center of the oscillation of the left plate and the right plate d [Eq. (5)].

In this paper, the computation is made for a single value of ϵ ($\epsilon = 0.1$), five values of a_w ($a_w = 0.01, 0.02, 0.05, 0.1, \text{ and } 0.5$), and various values of d in the range $d_0 \leq d \leq 3d_0$; here, $d_0 = 2\pi\sqrt{5/6} = 5.7357\cdots$, and its dimensional counterpart \tilde{d}_0 , i.e., $\tilde{d}_0 = d_0L = 2\pi(5R\tilde{T}_0/3)^{1/2}/\tilde{\omega}$, is the wavelength of the sinusoidal acoustic wave with angular frequency $\tilde{\omega}$ in an inviscid (Euler) gas.

In the present computation, we use uniform grids for x_1 and uniform time steps. The basic choices of the numerical parameters N and Δt are: N = 2000 ($d = d_0$) to 6000 ($d = 3d_0$) and $\Delta t \approx 1.9 \times 10^{-5}$ ($a_w = 0.01$ and 0.1) to 9.8×10^{-6} or 4.9×10^{-6} ($a_w = 0.5$). We also use coarser and finer grid systems for the accuracy tests. The details of the numerical parameters and the accuracy tests are summarized in Appendix C.

A. Profiles of macroscopic quantities

After the start of the oscillation of the left plate at t = 0, the plate sends out expansion and compression waves continuously. These waves interact with the waves reflected by the right



FIG. 2: Profiles of ρ , v_1 , T, and p in the time interval $199 < t/2\pi \le 200$ for $\epsilon = 0.1$, $a_w = 0.01$, and $d = d_0$ (hard-sphere molecules). (a) ρ , (b) v_1 , (c) T, and (d) p. The thin solid lines indicate the profiles at $t/2\pi = 199.1$, 199.2, ... 199.5, the dashed lines those at $t/2\pi = 199.6$, 199.7, ... 199.9, and the thick solid line the profile at $t/2\pi = 200$.

plate at rest and form a complicated flow field. But, the unsteady flow field tends to converge a time-periodic flow field after a few tens of oscillations of the plate.

In Figs. 2–6, we show the profiles of the physical quantities for hard-sphere molecules after the time-periodic state seems to have been established. Figures 2–4 show the profiles of the dimensionless density ρ , flow velocity v_1 , temperature T, and pressure p over a period 199 $< t/2\pi \leq 200$ (i.e., at $t/2\pi = 199.1$, 199.2, ..., 200) for $\epsilon = 0.1$ and $d = d_0$: Fig. 2 is for $a_w = 0.01$, Fig. 3 for $a_w = 0.1$, and Fig. 4 for $a_w = 0.5$. It should be noted that ρ , T, and p need corrections inside the Knudsen layer in order to give correct density, temperature, and pressure, as discussed in Appendix B 4. These corrections are omitted in Figs. 2–6 and Fig. 7 appearing later.

In the case of a small amplitude (and small Mach number) (Fig. 2), the profile of v_1 [Fig. 2(b)], which is fixed to be zero on the stationary plate, exhibits the structure similar to a standingwave with a node and two anti-nodes, and it is close to sinusoidal curves at $t/2\pi = 199.5$ and 200. The profiles of ρ , T, and p roughly show the two-node structure with a clearer nodelike point closer to the stationary plate ($x_1 \approx 4.4$). This standing-wave-like structure with node-like and anti-node-like points is less clear for a larger amplitude (Fig. 3) though v_1 still retains the structure [Fig. 3(b)]. For a large amplitude (Fig. 4), for which the nonlinearity becomes significant, the profiles deviate from the standing-wave-like profile, and no node-like or anti-node-like point is observed. In this case, ρ increases almost up to twice the initial density ($\rho = 1$) on the two plates, but decreases down to 60% of it in the middle of the gas. The Tbecomes almost twice the initial temperature (and the plate temperature) in the middle of the gas. There is a significant temperature jump on the resting plate because of the steep gradient of T and that of v_1 there [cf. Eq. (27b)].

Figures 5 and 6 show the profiles at $\epsilon = 0.1$ and $a_w = 0.1$ but for different d: $d = 1.2d_0$ (Fig. 5) and $2d_0$ (Fig. 6). For a slightly wider gap between the two plates (Fig. 5), the feature of the profile changes from that of Fig. 3. The standing-wave-like structure becomes less clear for v_1 but clearer for ρ , T, and p. In particular, ρ and T exhibit profiles with two node-like points and three anti-node-like points rather clearly [Figs. 5(a) and 5(c)]. On the other hand,



FIG. 3: Profiles of ρ , v_1 , T, and p in the time interval $199 < t/2\pi \le 200$ for $\epsilon = 0.1$, $a_w = 0.1$, and $d = d_0$ (hard-sphere molecules). (a) ρ , (b) v_1 , (c) T, and (d) p. See the caption of Fig. 2.



FIG. 4: Profiles of ρ , v_1 , T, and p in the time interval 199 $< t/2\pi \le 200$ for $\epsilon = 0.1$, $a_w = 0.5$, and $d = d_0$ (hard-sphere molecules). (a) ρ , (b) v_1 , (c) T, and (d) p. See the caption of Fig. 2.



FIG. 5: Profiles of ρ , v_1 , T, and p in the time interval $299 < t/2\pi \le 300$ for $\epsilon = 0.1$, $a_w = 0.1$, and $d = 1.2d_0$ (hard-sphere molecules). (a) ρ , (b) v_1 , (c) T, and (d) p. The thin solid lines indicate the profiles at $t/2\pi = 299.1$, 299.2, ... 299.5, the dashed lines those at $t/2\pi = 299.6$, 299.7, ... 299.9, and the thick solid line the profile at $t/2\pi = 300$.

for $d = 2d_0$ (Fig. 6), the standing-wave-like structure is clear (in particular, in the right half of the gas region) only for v_1 [Fig. 6(b)]. This difference between the profile of v_1 and those of ρ . T, and p is due to the difference in constraint on the plates. The v_1 takes the imposed values on the plate [cf. Eq. (27)], which corresponds to the fixed-end condition for a standing wave, whereas no constraint is imposed on ρ on the plates, which corresponds to the free-end condition for a standing wave. The condition for the temperature on the plates is somewhat intermediate, since the jump condition, Eq. (27), does not fix the values completely. Of course, the situation is not so simple because there is the viscous dissipation, and v_1 , ρ , and T are not independent but interact nonlinearly. Nevertheless, some behavior of the profiles may be understood from these properties. For instance, the oscillating plate inputs kinetic energy into the gas, so that the velocity field is affected by the plate most directly. In the cases of $d = d_0$ (Fig. 3) and $2d_0$ (Fig. 6), the standing-wave-like profile of v_1 is observed, as mentioned above. For $d = d_0$, the amplitude of v_1 in the gas exceeds the input amplitude $a_w = 0.1$ in the gas [Fig. 3(b)] in spite of the fact that the viscous dissipation should decay the amplitude. This amplification is caused by a sort of resonance (see the last paragraph in Sec. VIIB). It is also observed for $d = 2d_0$ [Fig. 6(b)]. In this case, because of the wider gas region, the waves attenuate more. Nevertheless, the amplitude of v_1 reaches almost the input value 0.1 in the gas. For $d = 1.2d_0$ [Fig. 5(b)], the amplitude of v_1 is smaller than the input value.

In Fig. 7, the profiles corresponding to Fig. 4 ($\epsilon = 0.1$, $a_w = 0.5$, and $d = d_0$) are shown for the BGK model. They are qualitatively the same as those in Fig. 4, but the maximum values of the density and temperature during one period is larger.

B. Momentum and energy transfer

Let us denote by \mathcal{P}^{L} and \mathcal{P}^{R} the x_{1} component of the momentum transferred from the left (oscillating) plate to the gas and that transferred from the gas to the right (resting) plate per



FIG. 6: Profiles of ρ , v_1 , T, and p in the time interval $399 < t/2\pi \le 400$ for $\epsilon = 0.1$, $a_w = 0.1$, and $d = 2d_0$ (hard-sphere molecules). (a) ρ , (b) v_1 , (c) T, and (d) p. The thin solid lines indicate the profiles at $t/2\pi = 399.1$, 399.2, ... 399.5, the dashed lines those at $t/2\pi = 399.6$, 399.7, ... 399.9, and the thick solid line the profile at $t/2\pi = 400$.

unit area and per unit time, respectively. We also denote by \mathcal{E}^{L} and \mathcal{E}^{R} the energy transferred from the left plate to the gas and that from the gas to the right plate, respectively. Then, if we neglect the terms of $O(\epsilon^2)$, they are expressed in the following form:

$$\mathcal{P}^{\mathbf{L}}(t) = \left(p - \frac{4}{3}\epsilon\Gamma_1 \frac{\partial v_1}{\partial x_1}\right)_{x_1 = x_w(t)},\tag{51a}$$

$$\mathcal{P}^{\mathrm{R}}(t) = \left(p - \frac{4}{3}\epsilon\Gamma_1 \frac{\partial v_1}{\partial x_1}\right)_{x_1=d},\tag{51b}$$

$$\mathcal{E}^{\mathbf{L}}(t) = \left[\left(p - \frac{4}{3} \epsilon \Gamma_1 \frac{\partial v_1}{\partial x_1} \right) v_w(t) - \frac{5}{4} \epsilon \Gamma_2 \frac{\partial T}{\partial x_1} \right]_{x_1 = x_w(t)},\tag{51c}$$

$$\mathcal{E}^{\mathrm{R}}(t) = \left(-\frac{5}{4}\epsilon\Gamma_2\frac{\partial T}{\partial x_1}\right)_{x_1=d}.$$
(51d)

These formulas are not subject to the Knudsen-layer corrections, as discussed in Appendix B 4. In addition, we introduce the time average of \mathcal{P}^{L} , \mathcal{P}^{R} , \mathcal{E}^{L} , and \mathcal{E}^{R} over one period from $t - 2\pi$ to t and denote them with an overline, i.e.,

$$\overline{\mathcal{W}}(t) = \frac{1}{2\pi} \int_{t-2\pi}^{t} \mathcal{W}(t') dt', \qquad (52)$$

where \mathcal{W} stands for \mathcal{P}^{L} , \mathcal{P}^{R} , \mathcal{E}^{L} , or \mathcal{E}^{R} .

Figures 8 and 9 show $(\overline{\mathcal{P}^{L}} - 1)/a_w$, $(\overline{\mathcal{P}^{R}} - 1)/a_w$, $\overline{\mathcal{E}^{L}}/a_w$, and $\overline{\mathcal{E}^{R}}/a_w$ at a long time, for which the time-periodic state seems to have been reached, versus d/d_0 at $\epsilon = 0.1$ for several values of a_w : Fig. 8 is for hard-sphere molecules, and Fig. 9 for the BGK model. The shown results are at $t/2\pi = 200$ for $1 \le d/d_0 \le 2$, at $t/2\pi = 400$ for $2 < d/d_0 \le 2.5$, and at $t/2\pi = 600$ for $2.5 < d/d_0 \le 3$. As is seen from the figures, $\overline{\mathcal{P}^{L}} = \overline{\mathcal{P}^{R}}$ and $\overline{\mathcal{E}^{L}} = \overline{\mathcal{E}^{R}}$ hold. This also indicates



FIG. 7: Profiles of ρ , v_1 , T, and p in the time interval $199 < t/2\pi \le 200$ for $\epsilon = 0.1$, $a_w = 0.5$, and $d = d_0$ (BGK model). (a) ρ , (b) v_1 , (c) T, and (d) p. The thin solid lines indicate the profiles at $t/2\pi = 199.1$, 199.2, ... 199.5, the dashed lines those at $t/2\pi = 199.6$, 199.7, ... 199.9, and the thick solid line the profile at $t/2\pi = 200$.

that the time-periodic state has been established. If \mathcal{P}^{L} , \mathcal{P}^{R} , \mathcal{E}^{L} , and \mathcal{E}^{R} make sinusoidal oscillation in time with period 2π as in the case of linear setting, $\overline{\mathcal{P}^{L}} - 1$, $\overline{\mathcal{P}^{R}} - 1$, $\overline{\mathcal{E}^{L}}$, and $\overline{\mathcal{E}^{R}}$ all vanish. In fact, when a_w is small ($a_w = 0.01$), these quantities are very small. For $a_w = 0.01$, $\overline{\mathcal{P}^{\mathrm{L}}} - 1$ and $\overline{\mathcal{P}^{\mathrm{R}}} - 1$ are negative for all $d/d_0 \ (\leq 3)$ in Fig. 8 and for $d/d_0 \lesssim 2$ in Fig. 9. This is because the average distance between the plates after long time is d, which is longer than the initial distance $d-a_w$. Therefore, the stationary plate is pulled (relative to the pressure exerted on the plate in the initial equilibrium state) by the gas, and the same is true for the oscillating plate. As d increases, the difference between d and $d - a_w$ becomes small, so that the pulling effect becomes small. Indeed, $\overline{\mathcal{P}^{\mathrm{L}}} - 1$ and $\overline{\mathcal{P}^{\mathrm{R}}} - 1$ are closer to zero for larger d/d_0 . When the amplitude (and the Mach number) a_w becomes large $(a_w = 0.01 \rightarrow 0.5)$, $\overline{\mathcal{P}^L} - 1$ and $\overline{\mathcal{P}^R} - 1$ become positive and larger, and $\overline{\mathcal{E}^{\mathrm{L}}}$ and $\overline{\mathcal{E}^{\mathrm{R}}}$ also increase. In other words, when a_w is not small, the stationary plate is pushed outward (i.e., in the positive x_1 direction), on the average, by the oscillating plate via the gas. The stationary plate also receives the energy (for all d), on the average, from the oscillating plate via the gas. The mechanism of this average momentum and energy transfer is explained in view of the interaction of the gas molecules with the moving plate in [21] (see Sec. 3.3 in [21]). Therefore, we do not repeat it here. If there is no stationary plate that receives the positive momentum relative to the initial pressure $(\overline{\mathcal{P}^R} - 1 > 0)$, the gas flows in the x_1 direction, i.e., toward infinity. This flow is called the acoustic stream [43, 44].

It is seen from Figs. 8 and 9 that both $\overline{\mathcal{P}^{\mathrm{L}}}$ (or $\overline{\mathcal{P}^{\mathrm{R}}}$) and $\overline{\mathcal{E}^{\mathrm{L}}}$ (or $\overline{\mathcal{E}^{\mathrm{R}}}$) exhibit local maxima and minima almost periodically with respect to d/d_0 . When a_w is small, the period is almost equal to $d_0/2$, that is, these quantities take local maxima near $d/d_0 = n/2$ with integer n and local minima near $d/d_0 = n/2 + 1/4$. For larger values of a_w ($a_w = 0.1 \rightarrow 0.5$), the period tends to increase.

The appearance of the local maxima at $d/d_0 \approx n/2$ for small a_w is reasonable for the following reason. Let us consider a column of an inviscid (or Euler) gas and a sinusoidal sound wave propagating along the column. If the column has reflective ends and its length is a multiple



FIG. 8: Average momentum and energy transfer over a period at the time-periodic state for $\epsilon = 0.1$ and some a_w (hard-sphere molecules). (a), (b) $(\overline{\mathcal{P}^{\mathrm{L}}} - 1)/a_w$ and $(\overline{\mathcal{P}^{\mathrm{R}}} - 1)/a_w$ vs d/d_0 , (c), (d) $\overline{\mathcal{E}^{\mathrm{L}}}/a_w$ and $\overline{\mathcal{E}^{\mathrm{R}}}/a_w$ vs d/d_0 . The shown results are at $t/2\pi = 200$ for $1 \leq d/d_0 \leq 2$, at $t/2\pi = 400$ for $2 < d/d_0 \leq 2.5$, and at $t/2\pi = 600$ for $2.5 < d/d_0 \leq 3$.

of $d_0/2 = \pi (5R\tilde{T}_0/3)^{1/2}/\tilde{\omega}$, the natural frequency of the column coincides with $\tilde{\omega}$. That is, a sinusoidal standing wave may be formed, and if one of the ends oscillates sinusoidally with angular frequency $\tilde{\omega}$, the amplitude of the standing wave is amplified. In other words, the resonance takes place. If there is no energy dissipation, the amplitude increases indefinitely. In a real gas, however, the energy dissipates and escapes from the column in the form of heat through the ends, so that the amplitude remains finite even in the case of resonance, i.e., even when $d/d_0 \approx n/2$. Nevertheless, the amplitude in these cases is still larger than that for other d. In other words, more energy and momentum are transmitted to the gas in this situation. This roughly explains the appearance of the local maxima at $d/d_0 \approx n/2$ when the amplitude of the oscillation of the plate a_w is small.

C. Approach to time-periodic state

As mentioned in Sec. VII A, the motion of the gas approaches a time-periodic state as time proceeds. In this section, we investigate the speed of approach to the time-periodic state. Here, we consider the case of $\epsilon = 0.1$, $a_w = 0.1$, and $d = d_0$ and regard the numerical solution for $199 < t/2\pi \le 200$ as the time-periodic solution. For any given time t, an integer n and a number $\tau \in (0, 2\pi]$ such that $t = (n-1) \cdot 2\pi + \tau$ are uniquely determined, i.e., n-1 is the maximum integer that is smaller than $t/2\pi$. Then, we define the difference between the solution at time t and the time-periodic solution by

$$h_{\text{diff}}(t, x_1) = h(t, x_1) - h(199 \cdot 2\pi + \tau, x_1), \tag{53}$$



FIG. 9: Average momentum and energy transfer over a period at the time-periodic state for $\epsilon = 0.1$ and some a_w (BGK model). (a), (b) $(\overline{\mathcal{P}^{\mathrm{L}}} - 1)/a_w$ and $(\overline{\mathcal{P}^{\mathrm{R}}} - 1)/a_w$ vs d/d_0 , (c), (d) $\overline{\mathcal{E}^{\mathrm{L}}}/a_w$ and $\overline{\mathcal{E}^{\mathrm{R}}}/a_w$ vs d/d_0 . The shown results are at $t/2\pi = 200$ for $1 \le d/d_0 \le 2$, at $t/2\pi = 400$ for $2 < d/d_0 \le 2.5$, and at $t/2\pi = 600$ for $2.5 < d/d_0 \le 3$.

where h represents the macroscopic quantities, i.e., $h = \rho$, v_1 , T, and p. Obviously, $h(199 \cdot 2\pi + \tau, x_1)$ ($0 < \tau \leq 2\pi$) indicates our time-periodic solution. If we plot $|h_{\text{diff}}|$ at a fixed x_1 as a function of t, it is oscillatory as is shown schematically in Fig. 10. Therefore, we pick up the peaks indicated by small circles in the figure. Each of these peaks corresponds to the maximum after the previous peak and can be picked up by the following procedure:

- 1. We first prepare the time-series data of all local maxima of $|h_{\text{diff}}|$.
- 2. We compare one of the local maxima with the next one. If the former maxima in smaller than the latter, we remove the former one from the time-series data.
- 3. We repeat the procedure 2 until there is no local maximum to be removed.

In Fig. 11, we show the curves joining the peaks thus obtained in semi-logarithmic scale for $x_1 = a_w = 0.1$ [Fig. 11(a)], $x_1 = (a_w + d)/2 = (0.1 + d_0)/2 = 2.9178 \cdots$ [Fig. 11(b)], and $x_1 = d = d_0 = 5.7357 \cdots$ [Fig. 11(c)] $[v_1$ is excluded in Fig. 11(c) because $v_1 = 0$ at $x_1 = d$]. For all ρ , v_1 , T, and p and at all three points, the curve tends to approach a straight line after the initial stage, and the gradients of all lines are almost the same for each of the hard-sphere molecules and the BGK model. If we denote the function corresponding to the curve by $F(t, x_1)$, then we have $|h_{\text{diff}}(t, x_1)| \leq F(t, x_1)$ and $\log_{10} F(t, x_1) \approx -\alpha' t + \beta'$ with α' and β' being constants. This is equivalent to write

$$|h_{\text{diff}}(t, x_1)| \le F(t, x_1) \approx C \exp(-\alpha t), \quad \text{for} \quad t \gg 1,$$
(54)

where C and α (= $\alpha' \ln 10 = \alpha' \times 2.302585 \cdots$) are constants, and α is seemingly independent of x_1 and the physical quantities but dependent on the molecular model. The approach to



FIG. 10: Schematic figure for $|h_{\text{diff}}|$ vs t. The small circle \circ indicates the peaks, each of which corresponds to the maximum after the previous peak.



FIG. 11: Curves joining the peaks obtained by the procedures 1-3 for $\epsilon = 0.1$, $a_w = 0.1$, and $d = d_0$. (a) at $x_1 = a_w = 0.1$, (b) $x_1 = (a_w + d)/2 = (0.1 + d_0)/2 = 2.9178 \cdots$, and (c) $x_1 = d = d_0 = 5.7357 \cdots$. The solid lines indicate $|\rho_{\text{diff}}|$, the dashed lines $|v_{1\text{diff}}|$, the dash-dotted lines $|T_{\text{diff}}|$, and the dash-double-dotted lines $|p_{\text{diff}}|$. The curve of $|v_{1\text{diff}}|$ is excluded in (c) because $v_1 = 0$ at $x_1 = d$.

the time-periodic state is faster for hard-sphere molecules. The curves of $F(t, x_1)$ cease to decrease when $F(t, x_1)$ becomes 10^{-5} to 10^{-7} . This is probably because the computation has reached the limit of accuracy. In fact, we have confirmed that the part for the exponential decay extends for finer space grids and time steps. We now pick up the interval of t in which each curve corresponding to $F(t, x_1)$ seems to be a straight line, that is, we take the interval $10 < t/2\pi \leq 30$ for hard-sphere molecules and $20 < t/2\pi \leq 50$ for the BGK model. Then, for this interval, we construct an approximate straight line, by the least square method, for

VIII. CONCLUDING REMARKS

In the present paper, we investigated the unsteady motion of a rarefied gas between two parallel plates caused when one of the plates starts a harmonic oscillation in its normal direction. We considered the case where the speed of oscillation of the plate is not necessarily small compared to the sonic speed (i.e., fully nonlinear setting), but the Knudsen number is small. Therefore, as an alternative to the Boltzmann or its model equations, for which accurate numerical computation for a long time is difficult for small Knudsen numbers, we decided to use the compressible Navier–Stokes equations and slip boundary conditions. However, it was practically impossible to find appropriate slip (or jump) boundary conditions for the compressible Navier–Stokes equations that can be applied immediately to the present problem. For example, one can find formulas for slip boundary conditions for compressible Navier–Stokes equations in [37]. However, the boundary is assumed to be at rest, and the numerical values of the coefficients in the formulas are not given. Therefore, we revisited this classical problem and derived the correct temperature jump condition for the present problem (Appendix B).

After we had prepared the boundary condition, we solved the compressible Navier–Stokes system numerically by a method suitable for the present problem containing a plate moving in its normal direction (Sec. VI). The result shows that the flow field approaches the time-periodic state after a few tens of oscillations of the plate. The properties of this time-periodic state were investigated in detail (Secs. VII B and VII C). When the speed of oscillation of the plate is not very small compared to the sonic speed, the momentum in the outward direction and the energy, averaged over a period, is transmitted to the stationary plate by the oscillating plate (Sec. VII B). This confirms the earlier result [21] based on the BGK model for the intermediate Knudsen numbers though the computation there was carried out until shorter times.

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Appendix A: The Boltzmann collision integral

1. Collision operators

In this appendix, we summarize the collision integral and related matters. The notation in the present paper is basically based on [25] (see Sec. 1.9 and Appendix A in [25]). However, since we avoided using to indicated dimensionless quantities in the main text, there are some differences in notations. For instance, [f, g, J(f, g)] in this appendix correspond to $[\hat{f}, \hat{g}, \hat{J}(\hat{f}, \hat{g})]$ in [25]. In addition, we occasionally use the bold face to indicate vectors and the dot \cdot the scalar product, e.g., $\boldsymbol{\zeta} = \zeta_i$ and $\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta}) = \alpha_j(\zeta_{*j} - \zeta_j)$, in Appendices A and B.

We start with the definition of the bilinear form J(f, g):

$$J(f, g) = \frac{1}{2} \int_{\boldsymbol{\zeta}_* \in \mathbf{R}^3, \, \boldsymbol{\alpha} \in \mathbf{S}^2} \left(f'g'_* + f'_*g' - fg_* - f_*g \right) \hat{B} \, d\Omega(\boldsymbol{\alpha}) \, d\boldsymbol{\zeta}_*.$$
(A1)

Here, we used the convention, i.e., $f = f(\zeta)$, $f_* = f(\zeta_*)$, $f' = f(\zeta')$, $f'_* = f(\zeta'_*)$, and the same for g; ζ , which was denoted by ζ_i in the main text, is the dimensionless molecular velocity; when a pair of molecules with velocities ζ and ζ_* collide, the velocities of the respective molecules after collision ζ' and ζ'_* are expressed as

$$\boldsymbol{\zeta}' = \boldsymbol{\zeta} + [\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta})] \boldsymbol{\alpha}, \qquad \boldsymbol{\zeta}'_* = \boldsymbol{\zeta}_* - [\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta})] \boldsymbol{\alpha}, \tag{A2}$$

where $\boldsymbol{\alpha}$ is the unit vector in the direction of $\boldsymbol{\zeta}' - \boldsymbol{\zeta}$; $d\boldsymbol{\zeta}_* = d\zeta_{*1}d\zeta_{*2}d\zeta_{*3}$ and $d\Omega(\boldsymbol{\alpha})$ is the solid-angle element around $\boldsymbol{\alpha}$; \hat{B} is a non-negative function of $|\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta})|/|\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|$ and $|\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|$, i.e.,

$$\hat{B} = \hat{B}\left(\frac{|\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta})|}{|\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|}, |\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|\right),\tag{A3}$$

which depends on the intermolecular potential, and $\hat{B} = |\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_* - \boldsymbol{\zeta})|/4\sqrt{2\pi}$ for hard-sphere molecules (see Sec. 1.9 and Appendix A in [25] for the details of \hat{B} , noting that \hat{B} here is the same as \hat{B} in [25]).

According to Sec. 1.9 in [25], the mean collision frequency $\bar{\nu}_{c0}$ and the mean free path l_0 in the equilibrium state at rest with density $\tilde{\rho}_0$ and temperature \tilde{T}_0 are, respectively, expressed as

$$\bar{\nu}_{c0} = (\tilde{\rho}_0/m)B_0, \qquad l_0 = (2/\sqrt{\pi})\tilde{c}_0/\bar{\nu}_{c0} = (2/\sqrt{\pi})\tilde{c}_0/(\tilde{\rho}_0/m)B_0,$$
 (A4)

where B_0 is defined by

$$B_{0} = \frac{1}{\tilde{\rho}_{0}^{2}} \int \tilde{f}_{0}(\tilde{\boldsymbol{\zeta}}) \tilde{f}_{0}(\tilde{\boldsymbol{\zeta}}_{*}) B\left(\frac{|\boldsymbol{\alpha} \cdot (\tilde{\boldsymbol{\zeta}}_{*} - \tilde{\boldsymbol{\zeta}})|}{|\tilde{\boldsymbol{\zeta}}_{*} - \tilde{\boldsymbol{\zeta}}|}, |\tilde{\boldsymbol{\zeta}}_{*} - \tilde{\boldsymbol{\zeta}}|\right) d\Omega(\boldsymbol{\alpha}) d\tilde{\boldsymbol{\zeta}} d\tilde{\boldsymbol{\zeta}}_{*}.$$
(A5)

Here, $\tilde{f}_0(\tilde{\boldsymbol{\zeta}})$ is the dimensional Maxwellian with density $\tilde{\rho}_0$, velocity 0, and temperature \tilde{T}_0 , i.e., $\tilde{f}_0(\tilde{\boldsymbol{\zeta}}) = \tilde{\rho}_0 (2\pi R \tilde{T}_0)^{-3/2} \exp(-|\tilde{\boldsymbol{\zeta}}|^2/2R \tilde{T}_0)$, and B is the dimensional counterpart of \hat{B} appearing in the dimensional form of the Boltzmann collision integral. In fact, \hat{B} is defined by $\hat{B} = B/B_0$. For hard-sphere molecules with diameter d_m ,

$$\bar{\nu}_{c0} = 2\sqrt{2\pi}\tilde{c}_0 d_m^2(\tilde{\rho}_0/m), \qquad l_0 = 1/\sqrt{2\pi} d_m^2(\tilde{\rho}_0/m).$$
 (A6)

When the intermolecular potential extends to infinity, the integral in Eq. (A5) generally diverges. See Sec. 1.9 in [25] for the treatment in such a case.

The linearized collision operator $\mathcal{L}(\cdot)$ is defined by

$$\mathcal{L}(\varphi) = 2J(E, E\varphi)/E$$

= $\int E(\zeta_*)(\varphi'_* + \varphi' - \varphi_* - \varphi) \hat{B}\left(\frac{|(\zeta_* - \zeta) \cdot \alpha|}{|\zeta_* - \zeta|}, |\zeta_* - \zeta|\right) d\Omega(\alpha) d\zeta_*,$ (A7)

where $\zeta_* = (\zeta_{*1}^2 + \zeta_{*2}^2 + \zeta_{*3}^2)^{1/2}$, $E(\cdot)$ is the function defined by Eq. (16b), and the same convention as in Eq. (A1) is used, i.e., $\varphi_* = \varphi(\zeta_*)$, $\varphi' = \varphi(\zeta')$, etc. Now we introduce the following extended linearized collision operator $\mathcal{L}_a(\cdot)$:

$$\mathcal{L}_{a}(\varphi) = \int E(\zeta_{*})(\varphi_{*}' + \varphi' - \varphi_{*} - \varphi) \hat{B}_{a}\left(\frac{|(\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}) \cdot \boldsymbol{\alpha}|}{|\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}|}, |\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}|\right) d\Omega(\boldsymbol{\alpha}) d\boldsymbol{\zeta}_{*}, \tag{A8}$$

where

$$\hat{B}_{a} = \hat{B}_{a} \left(\frac{|\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta})|}{|\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}|}, |\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}| \right) = \frac{1}{\sqrt{a}} \hat{B} \left(\frac{|\boldsymbol{\alpha} \cdot (\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta})|}{|\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}|}, \sqrt{a} |\boldsymbol{\zeta}_{*} - \boldsymbol{\zeta}| \right),$$
(A9)

and a is a positive quantity independent of $\boldsymbol{\zeta}$. Obviously, $\mathcal{L}_1(\varphi) = \mathcal{L}(\varphi)$ holds, and for hardsphere molecule, $\mathcal{L}_a(\varphi) = \mathcal{L}(\varphi)$ holds.

For the BGK model, the dimensionless collision integral J(f, f) in Eq. (8) is replaced by the following $J_{BGK}(f)$:

$$J_{\text{BGK}}(f) = \rho \left(f_e - f \right), \tag{A10}$$

where f_e is a local Maxwellian

$$f_e = \frac{\rho}{(\pi T)^{3/2}} \exp\left(-\frac{(\zeta_i - v_i)^2}{T}\right),$$
 (A11)

and ρ , v_i , and T are defined by Eq. (3a). In this model, the collision frequency ν_c of a molecule with velocity $\tilde{\zeta}_i$ depends neither on $\tilde{\zeta}_i$ nor on the shape of \tilde{f} and is assumed to be $A_c \tilde{\rho}$, where A_c is a constant. This corresponds to the Maxwell molecules in which the intermolecular potential is proportional to $1/r^4$ with r the distance between two molecules. Therefore, the mean collision frequency $\bar{\nu}_{c0}$ and the mean free path l_0 in the equilibrium state at rest with density $\tilde{\rho}_0$ and temperature \tilde{T}_0 become

$$\bar{\nu}_{c0} = A_c \tilde{\rho}_0, \qquad l_0 = (2/\sqrt{\pi})\tilde{c}_0/A_c \tilde{\rho}_0.$$
 (A12)

The linearized collision operators $\mathcal{L}_{aBGK}(\cdot)$ and $\mathcal{L}_{BGK}(\cdot)$ for the BGK model, which correspond to $\mathcal{L}_{a}(\cdot)$ and $\mathcal{L}(\cdot)$, respectively, take the following forms:

$$\sqrt{a}\mathcal{L}_{aBGK}(\varphi) = \int \left[1 + 2\boldsymbol{\zeta} \cdot \boldsymbol{\zeta}_* + \frac{2}{3}\left(\zeta^2 - \frac{3}{2}\right)\left(\zeta^2_* - \frac{3}{2}\right)\right]\varphi(\boldsymbol{\zeta}_*)E(\boldsymbol{\zeta}_*)d\boldsymbol{\zeta}_* - \varphi(\boldsymbol{\zeta}), \quad (A13a)$$
$$\mathcal{L}_{BGK}(\varphi) = \mathcal{L}_{1BGK}(\varphi), \quad (A13b)$$

$$\mathcal{L}_{BGK}(\varphi) = \mathcal{L}_{1BGK}(\varphi), \tag{A13b}$$

where $\zeta = (\zeta_1^2 + \zeta_2^2 + \zeta_3^2)^{1/2}$.

2. Functions \mathcal{A} and $\mathcal{B}^{(0)}$

The functions $\mathcal{A}(C, T)$ and $\mathcal{B}^{(0)}(C, T)$ occurring in the first-order term of the Chapman– Enskog solution (15) are defined as follows: $\mathcal{A}(\zeta, a)$ is the solution of the integral equation

$$\mathcal{L}_a[\zeta_i \mathcal{A}(\zeta, a)] = -\zeta_i \left(\zeta^2 - \frac{5}{2}\right),\tag{A14}$$

with the subsidiary condition

$$\int_0^\infty \zeta^4 \mathcal{A}(\zeta, a) E(\zeta) d\zeta = 0, \tag{A15}$$

and the function $\mathcal{B}^{(0)}(\zeta, a)$ is the solution of the integral equation

$$\mathcal{L}_{a}\left[\left(\zeta_{i}\zeta_{j}-\frac{1}{3}\zeta^{2}\delta_{ij}\right)\mathcal{B}^{(0)}(\zeta,a)\right] = -2\left(\zeta_{i}\zeta_{j}-\frac{1}{3}\zeta^{2}\delta_{ij}\right).$$
(A16)

In [25], $\mathcal{A}(\zeta, 1)$ is denoted by $A(\zeta)$, and $\mathcal{B}^{(0)}(\zeta, 1)$ by $B(\zeta)$:

$$\mathcal{A}(\zeta, 1) = A(\zeta), \qquad \mathcal{B}^{(0)}(\zeta, 1) = B(\zeta). \tag{A17}$$

Thus, for hard-sphere molecules, $\mathcal{A}(\zeta, a) = A(\zeta)$ and $\mathcal{B}^{(0)}(\zeta, a) = B(\zeta)$ for any a. The numerical values of $A(\zeta)$ and $B(\zeta)$ for hard-sphere molecules are tabulated in Table 3.1 of [25] (see also [45, 46]). These numerical data make it possible to perform numerical integration of the integrals in Eq. (19) and give Γ_1 and Γ_2 shown in Eq. (21a) [46].

For the BGK model, Eq. (A14) [with Eq. (A15)] and Eq. (A16), with $\mathcal{L}_a = \mathcal{L}_{aBGK}$, give the following solutions:

$$\mathcal{A}(\zeta, a) = \sqrt{a} \left(\zeta^2 - \frac{5}{2}\right), \qquad \mathcal{B}^{(0)}(\zeta, a) = 2\sqrt{a}, \tag{A18}$$

which give Γ_1 and Γ_2 shown in Eq. (21b).

In this appendix, we give an outline of the derivation of the jump boundary conditions (27).

1. Knudsen layer: Equation

We restrict ourselves to the derivation of the condition (27a) on the oscillating plate. From the result, the condition (27b) on the resting plate is derived immediately. Let us consider Eq. (26) near the moving plate $x_1 = x_w(t)$. Then, Φ is appreciable only in the thin layer of thickness of the order of ϵ adjacent to the plate. If we substitute Eq. (26) into the Boltzmann equation (8) and note the fact that

$$\frac{\partial f_{\rm CE}^{(1)}}{\partial t} + \zeta_1 \frac{\partial f_{\rm CE}^{(1)}}{\partial x_1} = \frac{1}{\epsilon} J(f_{\rm CE}^{(1)}, f_{\rm CE}^{(1)}) + O(\epsilon), \tag{B1}$$

we have

$$\epsilon \left(\frac{\partial f^{(0)}\Phi}{\partial t} + \zeta_1 \frac{\partial f^{(0)}\Phi}{\partial x_1}\right) = 2J(f_{\rm CE}^{(1)}, f^{(0)}\Phi) + \epsilon J(f^{(0)}\Phi, f^{(0)}\Phi) + O(\epsilon).$$
(B2)

Further, if we note that

$$\frac{\partial f^{(0)}\Phi}{\partial t} + \zeta_1 \frac{\partial f^{(0)}\Phi}{\partial x_1} = f^{(0)} \left(\frac{\partial \Phi}{\partial t} + \zeta_1 \frac{\partial \Phi}{\partial x_1} \right) + \Phi f^{(0)} \cdot O(1), \tag{B3a}$$

$$2J(f_{\rm CE}^{(1)}, f^{(0)}\Phi) = 2J(f^{(0)}, f^{(0)}\Phi) + \epsilon f^{(0)} \cdot O(\Phi),$$
(B3b)

Eq. (B2) is transformed to

$$\frac{\partial \Phi}{\partial t} + \zeta_1 \frac{\partial \Phi}{\partial x_1} = \frac{1}{\epsilon f^{(0)}} \cdot 2J(f^{(0)}, f^{(0)}\Phi) + O(\Phi).$$
(B4)

Let us introduce the new coordinate system $(\hat{t}, \eta, \hat{\zeta}_1, \hat{\zeta}_2, \hat{\zeta}_3)$ by

$$\hat{t} = t, \qquad \epsilon \eta = x_1 - x_w(t), \qquad \hat{\zeta}_1 = \zeta_1 - v_w(t), \qquad \hat{\zeta}_2 = \zeta_2, \qquad \hat{\zeta}_3 = \zeta_3,$$
(B5)

and change the variables from $(t, x_1, \zeta_1, \zeta_2, \zeta_3)$ to $(\hat{t}, \eta, \hat{\zeta}_1, \hat{\zeta}_2, \hat{\zeta}_3)$. More specifically, we write

$$\Phi(t, x_1, \zeta_1, \zeta_2, \zeta_3) = \hat{\Phi}(\hat{t}, \eta, \hat{\zeta}_1, \hat{\zeta}_2, \hat{\zeta}_3),$$
(B6)

and assume that

$$\frac{\partial \Phi}{\partial \eta} = O(\hat{\Phi}) \quad \text{and} \quad \hat{\Phi} \to 0 \text{ (rapidly) as } \eta \to \infty.$$
 (B7)

Then, LHS of (B4) becomes

$$LHS = \frac{\partial \hat{\Phi}}{\partial \hat{t}} + \frac{1}{\epsilon} \hat{\zeta}_1 \frac{\partial \hat{\Phi}}{\partial \eta} - \dot{v}_w(\hat{t}) \frac{\partial \hat{\Phi}}{\partial \hat{\zeta}_1}.$$
 (B8)

On the other hand, Eq. (A1) gives, with the same convention,

$$2J(f^{(0)}, f^{(0)}\Phi)(f^{(0)})^{-1} = \int f_*^{(0)}(\Phi'_* + \Phi' - \Phi_* - \Phi) \hat{B}\left(\frac{|(\boldsymbol{\zeta}_* - \boldsymbol{\zeta}) \cdot \boldsymbol{\alpha}|}{|\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|}, |\boldsymbol{\zeta}_* - \boldsymbol{\zeta}|\right) d\Omega(\boldsymbol{\alpha}) d\boldsymbol{\zeta}_*,$$
(B9)

where

$$f_*^{(0)} = \frac{\rho}{(\pi T)^{3/2}} \exp\left(-\frac{(\zeta_{*1} - v_1)^2 + \zeta_{*2}^2 + \zeta_{*3}^2}{T}\right),\tag{B10}$$

and the fact that $f^{(0)'}f_*^{(0)'} = f^{(0)}f_*^{(0)}$ has been used. We consider the range of η for which Φ is appreciable, i.e., $\eta = O(1)$. Then, for $h = \rho$, v_1 , and T, we can write

$$h = h(\hat{t}, x_w(\hat{t}) + \epsilon \eta) = h_B(\hat{t}) + O(\epsilon \eta), \tag{B11}$$

where $h_B(\hat{t}) \equiv h(\hat{t}, x_w(\hat{t}))$ indicates the value on the plate. If we assume Eq. (25), then from Eq. (B11) we have

$$\rho = \rho_B(\hat{t}) + O(\epsilon\eta), \qquad v_1 = v_w(\hat{t}) + O(\epsilon(\eta+1)), \qquad T = 1 + O(\epsilon(\eta+1)).$$
(B12)

Therefore, $f_*^{(0)}$ can be expressed as

$$f_*^{(0)} = \frac{\rho_B}{\pi^{3/2}} \exp\left(-\left[(\zeta_{*1} - v_w)^2 + \zeta_{*2}^2 + \zeta_{*3}^2\right]\right) \left[1 + O(\epsilon(\eta + 1))\right]$$

= $\rho_B E(\hat{\zeta}_*) \left[1 + O(\epsilon(\eta + 1))\right].$ (B13)

In the second line, the new variable $\hat{\zeta}_i$ [Eq. (B5)] and $E(\cdot)$ defined in Eq. (16b) are used; more specifically, $\hat{\zeta}_{*i}$ is defined by Eq. (B5) with $\zeta_i = \zeta_{*i}$, and $\hat{\zeta}_* = (\hat{\zeta}_{*1}^2 + \hat{\zeta}_{*2}^2 + \hat{\zeta}_{*3}^2)^{1/2}$. With this expression, Eq. (B9) can be expressed in terms of the new variables as

$$2J(f^{(0)}, f^{(0)}\Phi)(f^{(0)})^{-1} = \rho_B \int E(\hat{\zeta}_*)(\hat{\Phi}'_* + \hat{\Phi}' - \hat{\Phi}_* - \hat{\Phi}) \\ \times \hat{B}\left(\frac{|(\hat{\zeta}_* - \hat{\zeta}) \cdot \alpha|}{|\hat{\zeta}_* - \hat{\zeta}|}, |\hat{\zeta}_* - \hat{\zeta}|\right) d\Omega(\alpha) d\hat{\zeta}_* + O(\epsilon(\eta + 1)\hat{\Phi}),$$
(B14)

where $\hat{\zeta}'_i$ and $\hat{\zeta}'_{*i}$ are defined by Eq. (B5) with $\zeta_i = \zeta'_i$ and $\zeta_i = \zeta'_{*i}$, respectively, so that $\hat{\zeta}_i$, $\hat{\zeta}_{*i}$, $\hat{\zeta}'_{i}$, and $\hat{\zeta}'_{*i}$ satisfy the same relations as Eq. (A2). In Eq. (B14), the same convention is used: More specifically, $\hat{\Phi} = \hat{\Phi}(\hat{t}, \eta, \hat{\zeta}_i)$, $\hat{\Phi}_* = \hat{\Phi}(\hat{t}, \eta, \hat{\zeta}_{*i})$, $\hat{\Phi}' = \hat{\Phi}(\hat{t}, \eta, \hat{\zeta}'_i)$, and $\hat{\Phi}'_* = \hat{\Phi}(\hat{t}, \eta, \hat{\zeta}'_{*i})$. In summery, Eqs. (B8) and (B14) lead to the following expression of Eq. (B4):

$$\hat{\zeta}_1 \frac{\partial \hat{\Phi}}{\partial \eta} = \rho_B \mathcal{L}(\hat{\Phi}) + O(\epsilon(\eta + 1)\hat{\Phi}), \tag{B15}$$

where the linearized collision operator defined by Eq. (A7) is used, i.e.,

$$\mathcal{L}(\hat{\Phi}) = \int E(\hat{\zeta}_*)(\hat{\Phi}'_* + \hat{\Phi}' - \hat{\Phi}_* - \hat{\Phi}) \hat{B}\left(\frac{|(\hat{\zeta}_* - \hat{\zeta}) \cdot \boldsymbol{\alpha}|}{|\hat{\zeta}_* - \hat{\zeta}|}, |\hat{\zeta}_* - \hat{\zeta}|\right) d\Omega(\boldsymbol{\alpha}) d\hat{\zeta}_*.$$
 (B16)

Since $\hat{\Phi}$ vanishes rapidly as $\eta \to \infty$, the term of $O(\epsilon(\eta + 1)\hat{\Phi})$ in Eq. (B15), which is of $O(\epsilon)$ for finite η , vanish rapidly as $\eta \to \infty$. If we neglect this term, introduce a new coordinate y by

$$y = \rho_B \eta, \tag{B17}$$

and denote

$$\hat{\Phi}(\hat{t},\,\eta,\,\hat{\zeta}_i) = \phi(\hat{t},\,y,\,\hat{\zeta}_i),\tag{B18}$$

then, we have the following equation for ϕ :

$$\hat{\zeta}_1 \frac{\partial \phi}{\partial y} = \mathcal{L}(\phi) \qquad (0 < y < \infty).$$
 (B19)

2. Knudsen layer: Boundary condition

Next, we derive the boundary conditions for Eq. (B19). The velocity distribution function of the form of Eq. (26) has to satisfy the boundary condition on the plate, Eq. (10a), that is,

$$f_B^{(0)}(1 + \Psi_B \epsilon + \Phi_B \epsilon) = \frac{\sigma_w}{\pi^{3/2}} \exp\left(-\left\{ [\zeta_1 - v_w(t)]^2 + \zeta_2^2 + \zeta_3^2 \right\} \right),$$

for $\zeta_1 - v_w(t) > 0,$ (B20a)

$$\sigma_w = -2\pi^{1/2} \int_{\zeta_1 - v_w(t) < 0} [\zeta_1 - v_w(t)] f_B^{(0)} (1 + \Psi_B \epsilon + \Phi_B \epsilon) d\boldsymbol{\zeta}, \tag{B20b}$$

where the subscript B indicates the value on the plate, i.e., at $x_1 = x_w(t)$ or $\eta = 0$, and the terms of $O(\epsilon^2)$ are omitted.

Now we recall Eq. (25a), i.e., $v_{1B} - v_w(t) = O(\epsilon)$ and $T_B - 1 = O(\epsilon)$ and let

$$v_{1B} - v_w(t) = \check{v}_1 \epsilon, \qquad T_B - 1 = \check{T} \epsilon.$$
 (B21)

Then, $\check{v}_1 = O(1)$ and $\check{T} = O(1)$. By expanding v_{1B} around $v_w(t)$ and T_B around 1 in $f_B^{(0)}$ and Ψ_B , the boundary value $f_B^{(0)}(1 + \Psi_B \epsilon)$ is expressed in the following form:

$$f_B^{(0)}(1+\Psi_B\epsilon) = \rho_B E(\hat{\zeta}) \left[1 + 2\hat{\zeta}_1 \check{v}_1 \epsilon + \left(\hat{\zeta}^2 - \frac{3}{2}\right) \check{T}\epsilon - \frac{1}{\rho_B} \left(\hat{\zeta}_1^2 - \frac{1}{3}\hat{\zeta}^2\right) \left(\frac{\partial v_1}{\partial x_1}\right)_B \mathcal{B}^{(0)}(\hat{\zeta}, 1) \epsilon - \frac{1}{\rho_B}\hat{\zeta}_1 \left(\frac{\partial T}{\partial x_1}\right)_B \mathcal{A}(\hat{\zeta}, 1) \epsilon + O(\epsilon^2) \right],$$
(B22)

where the variable $\hat{\zeta}_i$ [Eq. (B5)] and $E(\cdot)$ [Eq. (16b)] are used, and $\hat{\zeta} = (\hat{\zeta}_1^2 + \hat{\zeta}_2^2 + \hat{\zeta}_3^2)^{1/2}$. If we use this expression in the integrand in Eq. (B20b) and carry out possible integrations, we have the following expression for σ_w :

$$\sigma_w = \rho_B [1 + \delta \epsilon + O(\epsilon^2)], \qquad (B23a)$$

$$\delta = -\sqrt{\pi} \check{v}_1 + \frac{1}{2} \check{T} - \frac{1}{3\rho_B} \left(\frac{\partial v_1}{\partial x_1}\right)_B \int_0^\infty r^5 \mathcal{B}^{(0)}(r, 1) e^{-r^2} dr$$

$$- 2\sqrt{\pi} \int_{\hat{\zeta}_1 < 0} \hat{\zeta}_1 \Phi_B E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}. \qquad (B23b)$$

Let us note that

$$\Phi_B = \Phi(t, x_w(t), \zeta_i) = \hat{\Phi}(\hat{t}, \eta = 0, \hat{\zeta}_i) = \phi(\hat{t}, y = 0, \hat{\zeta}_i).$$
(B24)

Then, Eq. (B20), together with Eqs. (B22)–(B24), gives the boundary condition for ϕ on the plate, that is,

Here, the terms of $O(\epsilon)$ have been neglected, and $\mathcal{A}(\hat{\zeta}, 1)$ and $\mathcal{B}^{(0)}(\hat{\zeta}, 1)$ have been replaced by $A(\hat{\zeta})$ and $B(\hat{\zeta})$ [cf. Eq. (A17)]. In addition, because of Eq. (B7), ϕ should vanish rapidly as $y \to \infty$, i.e.,

$$\phi \to 0$$
 as $y \to \infty$. (B26)

Equations (B25) and (B26) form the boundary conditions for Eq. (B19).

3. Jump boundary conditions

Before analyzing Eqs. (B19), (B25), and (B26), we integrate Eq. (B19) multiplied by $E(\hat{\zeta})$ with respect to $\hat{\zeta}_i$ over its whole space. Then, we have

$$\frac{\partial}{\partial y} \int \hat{\zeta}_1 \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}} = 0, \tag{B27}$$

since the integral of $\mathcal{L}(\phi)$ vanishes. Because of Eq. (B26),

$$\int \hat{\zeta}_1 \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}} = 0 \tag{B28}$$

holds for any y. Therefore, we have

$$\int \hat{\zeta}_1 \phi(\hat{t}, y = 0, \hat{\zeta}_i) E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}} = 0.$$
(B29)

On the other hand, the boundary condition (10a) satisfies the impermeability condition,

$$\int [\zeta_1 - v_w(t)] f_B d\boldsymbol{\zeta} = \int \hat{\zeta}_1 f_B^{(0)} [1 + \Psi_B \epsilon + \Phi_B \epsilon + O(\epsilon^2)] d\hat{\boldsymbol{\zeta}} = 0.$$
(B30)

Since $f_B^{(0)} = \rho_B E(\hat{\zeta})[1 + O(\epsilon)]$ and $\Phi_B = \phi(\hat{t}, y = 0, \hat{\zeta}_i)$ [Eq. (B24)], Eqs. (B29) and (B30) give the relation

$$\int \hat{\zeta}_1 f_B^{(0)}(1+\Psi_B \epsilon) d\hat{\boldsymbol{\zeta}} = O(\epsilon^2).$$
(B31)

Using Eq. (B22), performing the integration, and noting that $\int \hat{\zeta}_1^2 \mathcal{A}(\hat{\zeta}, 1) E(\hat{\zeta}) d\hat{\zeta} = 0$, we obtain the following relation:

$$\check{v}_1 = [v_{1B} - v_w(t)]/\epsilon = O(\epsilon).$$
(B32)

Therefore, we can neglect \check{v}_1 in Eq. (B25).

The mathematical structure of the solution of the half-space boundary-value problem, Eqs. (B19), (B25), and (B26), is discussed in [25] on the basis of the theorem conjectured in [47], proved in [48], and further analyzed in [49] (the reader is also referred to the more recent overview [50]). In the present case, where $\check{v}_1 = 0$ in Eq. (B25), the unique solution exists when \check{T} is related to $(1/\rho_B)(\partial v_1/\partial x_1)_B$ and $(1/\rho_B)(\partial T/\partial x_1)_B$ appropriately. This relation, which we will obtain below, gives the boundary condition for the temperature.

Let us set

$$\phi = \frac{1}{\rho_B} \left(\frac{\partial v_1}{\partial x_1} \right)_B \phi_v + \frac{1}{\rho_B} \left(\frac{\partial T}{\partial x_1} \right)_B \phi_T, \tag{B33a}$$

$$\check{T} = \frac{1}{\rho_B} \left(\frac{\partial v_1}{\partial x_1} \right)_B \alpha_v + \frac{1}{\rho_B} \left(\frac{\partial T}{\partial x_1} \right)_B \alpha_T,$$
(B33b)

and insert them in Eqs. (B19), (B25) (with $\check{v}_1 = 0$), and (B26). Then, because of the linearity of the problem, we obtain the equations and boundary conditions for ϕ_v and ϕ_T : For ϕ_v , the equation and boundary conditions are

$$\hat{\zeta}_1 \frac{\partial \phi_v(\hat{t}, y, \hat{\boldsymbol{\zeta}})}{\partial y} = \mathcal{L}[\phi_v(\hat{t}, y, \hat{\boldsymbol{\zeta}})], \tag{B34a}$$
$$\phi_v(\hat{t}, 0, \hat{\boldsymbol{\zeta}}) = -2\sqrt{\pi} \int_{-\infty}^{\infty} \hat{\zeta}_{*1} \phi_v(\hat{t}, 0, \hat{\boldsymbol{\zeta}}_*) E(\hat{\zeta}_*) d\hat{\boldsymbol{\zeta}}_* - \alpha_v\left(\hat{\zeta}^2 - 2\right)$$

$$(t, 0, \boldsymbol{\zeta}) = -2\sqrt{\pi} \int_{\hat{\zeta}_{*1} < 0} \zeta_{*1} \phi_v(t, 0, \boldsymbol{\zeta}_*) E(\zeta_*) d\boldsymbol{\zeta}_* - \alpha_v \left(\zeta^2 - 2\right) - \frac{1}{3} \int_0^\infty r^5 B(r) e^{-r^2} dr + \left(\hat{\zeta}_1^2 - \frac{1}{3}\hat{\zeta}^2\right) B(\hat{\zeta}), \qquad (\hat{\zeta}_1 > 0),$$
(B34b)

$$\phi_v(\hat{t}, y, \hat{\zeta}) \to 0, \qquad (y \to \infty),$$
(B34c)

and for ϕ_T , they are

$$\hat{\zeta}_{1} \frac{\partial \phi_{T}(\hat{t}, y, \hat{\boldsymbol{\zeta}})}{\partial y} = \mathcal{L}[\phi_{T}(\hat{t}, y, \hat{\boldsymbol{\zeta}})],$$
(B35a)
$$\phi_{T}(\hat{t}, 0, \hat{\boldsymbol{\zeta}}) = -2\sqrt{\pi} \int_{\hat{\zeta}_{*1} < 0} \hat{\zeta}_{*1} \phi_{T}(\hat{t}, 0, \hat{\boldsymbol{\zeta}}_{*}) E(\hat{\zeta}_{*}) d\hat{\boldsymbol{\zeta}}_{*} - \alpha_{T} \left(\hat{\zeta}^{2} - 2\right) + \hat{\zeta}_{1} A(\hat{\zeta}),$$
(B35b)

$$\phi_T(\hat{t}, y, \hat{\boldsymbol{\zeta}}) \to 0, \qquad (y \to \infty).$$
 (B35c)

The undetermined constants α_v and α_T are determined together with the solutions ϕ_v and ϕ_T , respectively.

• Problem for ϕ_T

The problem for ϕ_T , Eq. (B35), is exactly the same as the classical temperature jump problem (see, e.g., [27, 51–54]). For hard-sphere molecules, the problem was solved numerically in [54] (see also [46]). We note that $(y, \hat{\zeta}_i, \phi_T, \alpha_T)$ (here) = $(x_1, \zeta_i, \Phi_K, \beta)$ (in [54]) = $(\eta, \zeta_i, \Phi_1, b)$ (in [46]) and α_T (here) = d_1 (in [25]; see Sec. 3.1.5 in [25]). For the BGK model, an accurate value of α_T was obtained in [55, 56]: α_T (here) = d_1 (in [55, 56]). In summary, we have from [25] the values listed in Eq. (28), i.e.,

$$\alpha_T = 2.4001$$
 (hard-sphere molecules), $\alpha_T = 1.30272$ (BGK model), (B36)

for the diffuse reflection.

• Problem for ϕ_v

The problem for ϕ_v , Eq. (B34), has been studied rarely because it does not correspond to a specific half-space problem of physical interest. It has appeared only in the generalized slip flow theory (linear theory) with evaporation and condensation on the boundary [25, 56] (BGK model) and that for time-dependent problems with solid boundary [34–36] (hard-sphere molecules). For instance, $(y, \hat{\zeta}_1, \phi_v, \alpha_v)$ (here) = $(\eta, \mu\zeta, \phi_5, c_5^{(0)})$ (in [35]), and α_v (here) = $(4/3)d_6$ in [25, 56]. To summarize, we have, from [35] for hard-sphere molecules and from [25] for the BGK model, the values of α_v listed in Eq. (28), i.e.,

$$\alpha_v = 0.45957$$
 (hard-sphere molecules), $\alpha_v = 0.44045$ (BGK model), (B37)

for the diffuse reflection.

In summary, Eqs. (B21), (B32), and (B33b) give v_{1B} and T_B in the following form:

$$v_{1B} - v_w(t) = 0, \qquad T_B - 1 = \frac{1}{\rho_B} \left(\frac{\partial v_1}{\partial x_1}\right)_B \alpha_v \epsilon + \frac{1}{\rho_B} \left(\frac{\partial T}{\partial x_1}\right)_B \alpha_T \epsilon,$$
 (B38)

where the term of $O(\epsilon^2)$ in Eq. (B21) has been neglected. With the values of α_v and α_T mentioned above, these relations give the boundary conditions for the Navier–Stokes equations on the oscillating plate, i.e., at $x_1 = x_w(t)$ [see Eq. (27a)]. The boundary conditions on the plate at rest, i.e., at $x_1 = d$, is obtained by replacing (x_1, v_1, v_w) with $(d - x_1, -v_1, 0)$ in Eq. (B38) [see Eq. (27b)] and noting that the subscript *B* indicates the values at $x_1 = d$.

As mentioned in the main text, there are some problems in the existing slip boundary conditions for the compressible Navier–Stokes equations, and incorrect formulas are often used. Therefore, it is worth deriving the correct formulas in the general setting, i.e., for boundaries with arbitrary shape, velocities, and temperatures. This will be the subject of our forthcoming paper [57].

4. Macroscopic quantities inside the Knudsen layer

Let h stand for any of the macroscopic quantities ρ , v_i , T, p, p_{ij} , and q_i in this subsection. The general relation between the velocity distribution function f and h is given by Eqs. (3a) and (3b). For convenience in the following discussions, we use subscript "CE" to indicate h associated with the Chapman–Enskog solution $f_{CE}^{(1)}$. To be more specific, h_{CE} (or ρ_{CE} , v_{CEi} , T_{CE} , p_{CE} , p_{CEij} , and q_{CEi}) in this subsection = h (or ρ , v_i , T, p, p_{ij} , and q_i) appeared in Secs. IV–VII and in Appendices B 1–B 3. The Chapman–Enskog solution is subject to the correction inside the Knudsen layer [cf. Eq. (26)] to provide the correct solution there. We rewrite Eq. (26) in the following form using the subscript "tot" to indicate the correct solution inside the Knudsen layer:

$$f_{\text{tot}} = f_{CE}^{(1)} + f^{(0)}\Phi + O(\epsilon^2) = f^{(0)}(1 + \Psi\epsilon + \Phi\epsilon) + O(\epsilon^2),$$
(B39)

where it is noted that the macroscopic quantities contained in $f^{(0)}$ are h_{CE} . Now we denote the correct macroscopic quantities inside the Knudsen layer by h_{tot} and express it as

$$h_{\rm tot} = h_{\rm CE} + h_{\rm K}^{(1)} \epsilon + O(\epsilon^2).$$
 (B40)

In other words, this provides the definition of $h_{\rm K}^{(1)}$.

We derive $h_{\rm K}^{(1)}$ in the Knudsen layer at the oscillating plate $[x_1 = x_w(t)]$. The corresponding result in the Knudsen-layer at the resting plate $(x_1 = d)$ can be obtained immediately. We substitute Eqs. (B39) and (B40) into the general relations, Eqs. (3a) and (3b) (with $f = f_{\rm tot}$, and $h = h_{\rm tot}$) and take into account that $f_{\rm CE}^{(1)}$ and $h_{\rm CE}$ also satisfy Eqs. (3a) and (3b). Then, we simplify the expressions by using Eq. (B12) for $h_{\rm CE}$ contained in the factors multiplied by Φ and neglect the terms of $O(\epsilon^2)$. After changing the integration variables from ζ_i to $\hat{\zeta}_i$ (thus changing from Φ to ϕ) [cf. Eqs. (B5), (B6), and (B18)], we obtain the following expressions of $h_{\rm K}^{(1)}$:

$$\rho_{\rm K}^{(1)} = (\rho_{\rm CE})_B \int \phi E(\hat{\zeta}) d\hat{\zeta}, \qquad v_{\rm Ki}^{(1)} = \int \hat{\zeta}_i \phi E(\hat{\zeta}) d\hat{\zeta}, \tag{B41a}$$

$$T_{\rm K}^{(1)} = \frac{2}{3} \int \left(\hat{\zeta}_j^2 - \frac{3}{2}\right) \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}, \qquad p_{\rm K}^{(1)} = (\rho_{\rm CE})_B T_{\rm K}^{(1)} + \rho_{\rm K}^{(1)}, \tag{B41b}$$

$$p_{\mathrm{K}ij}^{(1)} = 2(\rho_{\mathrm{CE}})_B \int \hat{\zeta}_i \hat{\zeta}_j \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}, \qquad q_{\mathrm{K}i}^{(1)} = (\rho_{\mathrm{CE}})_B \int \hat{\zeta}_i \left(\hat{\zeta}_j^2 - \frac{5}{2}\right) \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}. \tag{B41c}$$

From the symmetry of the problem [cf. Eq. (4)], $v_{\text{K2}}^{(1)} = v_{\text{K3}}^{(1)} = 0$, $p_{\text{K12}}^{(1)} = p_{\text{K21}}^{(1)} = p_{\text{K23}}^{(1)} = p_{\text{K32}}^{(1)} = 0$. Integrating Eq. (B19) multiplied by $\hat{\zeta}_1 E(\hat{\zeta})$ and that multiplied by $(\hat{\zeta}_j^2 - 5/2)E(\hat{\zeta})$ with respect to $\hat{\zeta}_i$ over its whole space and repeating the argument that led to Eq. (B28), we observe that $p_{\text{K11}}^{(1)} = q_{\text{K11}}^{(1)} = 0$.

whole space and repeating the argument that led to Eq. (B28), we observe that $p_{K11}^{(1)} = q_{K1}^{(1)} = 0$. Now we remove the subscript "CE" and summarize the correct (nontrivial) macroscopic quantities h_{tot} inside the Knudsen layer. That is, if we neglect the terms of $O(\epsilon^2)$, we have

$$\rho_{\text{tot}} = \rho + \epsilon \rho_{\text{K}}^{(1)}, \qquad v_{\text{tot}1} = v_1, \qquad T_{\text{tot}} = T + \epsilon T_{\text{K}}^{(1)}, \qquad p_{\text{tot}} = p + \epsilon p_{\text{K}}^{(1)}, \qquad (B42a)$$

$$p_{\text{tot}11} = p - \frac{4}{3}\epsilon\Gamma_1(T)\frac{\partial v_1}{\partial x_1}, \qquad p_{\text{tot}22} = p + \epsilon p_{\text{K}22}^{(1)}, \qquad p_{\text{tot}33} = p + \epsilon p_{\text{K}33}^{(1)},$$
(B42b)

$$q_{\text{tot1}} = -\frac{5}{4}\epsilon\Gamma_2(T)\frac{\partial T}{\partial x_1},\tag{B42c}$$

where

$$\rho_{\rm K}^{(1)} = \rho_B \int \phi E(\hat{\zeta}) d\hat{\zeta}, \qquad T_{\rm K}^{(1)} = \frac{2}{3} \int \left(\hat{\zeta}_j^2 - \frac{3}{2}\right) \phi E(\hat{\zeta}) d\hat{\zeta}, \tag{B43a}$$

$$p_{\rm K}^{(1)} = \rho_B T_{\rm K}^{(1)} + \rho_{\rm K}^{(1)}, \qquad p_{{\rm K}ij}^{(1)} = 2\rho_B \int \hat{\zeta}_i \hat{\zeta}_j \phi E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}.$$
 (B43b)

If we insert Eq. (B33a) in Eq. (B43a), we have

$$\rho_{\rm K}^{(1)} = \left(\frac{\partial v_1}{\partial x_1}\right)_B \Omega_v(y) + \left(\frac{\partial T}{\partial x_1}\right)_B \Omega_T(y),\tag{B44a}$$

$$T_{\rm K}^{(1)} = \frac{1}{\rho_B} \left(\frac{\partial v_1}{\partial x_1}\right)_B \Theta_v(y) + \frac{1}{\rho_B} \left(\frac{\partial T}{\partial x_1}\right)_B \Theta_T(y),\tag{B44b}$$

where

$$\Omega_{v}(y) = \int \phi_{v} E(\hat{\zeta}) d\hat{\zeta}, \qquad \Omega_{T}(y) = \int \phi_{T} E(\hat{\zeta}) d\hat{\zeta}, \tag{B45a}$$

$$\Theta_v(y) = \frac{2}{3} \int \left(\hat{\zeta}_j^2 - \frac{3}{2}\right) \phi_v E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}, \qquad \Theta_T(y) = \frac{2}{3} \int \left(\hat{\zeta}_j^2 - \frac{3}{2}\right) \phi_T E(\hat{\zeta}) d\hat{\boldsymbol{\zeta}}. \tag{B45b}$$

For hard-sphere molecules, $\Omega_T(y)$ and $\Theta_T(y)$ were obtained in [54] and [46]. More precisely, $[y, \Omega_T(y), \Theta_T(y)]$ (here) = $[x_1, \Omega(x_1), \Theta(x_1)]$ (in [54]) = $[\eta, \Omega(\eta), \Theta(\eta)]$ (in [46]) = $[\eta, \Omega_1(\eta), \Theta_1(\eta)]$ (in [25]). These functions are tabulated in the respective references; for example, see Table 3.2 on page 86 of [25]. On the other hand, $\Omega_v(y)$ and $\Theta_v(y)$ were obtained in [35]; $[y, \Omega_v(y), \Theta_v(y)]$ (here) = $[\eta, \Omega_5^{(0)}(\eta), \Theta_5^{(0)}(\eta)]$ (in [35]), and $\Omega_5^{(0)}(\eta)$ and $\Theta_5^{(0)}(\eta)$ are tabulated in Table 5.2 of [35].

For the BGK model, $\Omega_T(y)$ and $\Theta_T(y)$ were obtained, e.g., in [56]; $[y, \Omega_T(y), \Theta_T(y)]$ (here) = $[\eta, \Omega_1(\eta), \Theta_1(\eta)]$ (in [56]) = $[\eta, \Omega_1(\eta), \Theta_1(\eta)]$ (in [25]), and these functions are tabulated in the respective references; for example, see Table 3.3 on page 87 of [25]. As for $\Omega_v(y)$ and $\Theta_v(y)$, they were also obtained in [56]; $[y, \Omega_v(y), \Theta_v(y)]$ (here) = $[\eta, (4/3)\Omega_6(\eta), (4/3)\Theta_6(\eta)]$ (in [56]) = $[\eta, (4/3)\Omega_6(\eta), (4/3)\Theta_6(\eta)]$ (in [25]), and $\Omega_6(\eta)$ and $\Theta_6(\eta)$ are tabulated, for instance, in Table 3.3 on page 87 of [25].

Once the data of $\Omega_v(y)$, $\Theta_v(y)$, $\Omega_T(y)$, and $\Theta_T(y)$ are obtained from the references quoted above, the correct macroscopic quantities ρ_{tot} , T_{tot} , and p_{tot} are obtained from Eqs. (B42a), (B43b), and (B44) (we omit the result of $p_{\text{K22}}^{(1)}$ and $p_{\text{K33}}^{(1)}$). We emphasize that v_{tot1} , p_{tot11} , and q_{tot1} are free from the Knudsen-layer corrections.

The form of Eq. (B42) is the same in the Knudsen layer at the resting plate $(x_1 = d)$. The corresponding corrections $\rho_{\rm K}^{(1)}$ and $T_{\rm K}^{(1)}$ are obtained by replacing (x_1, v_1) with $(d - x_1, -v_1)$, defining y by $y = \rho_B(d - x_1)/\epsilon$, and noting that the subscript B indicates the values at $x_1 = d$ in Eq. (B44).

Appendix C: Data and accuracy test for numerical computation

First we summarize the data for the numerical computation. We have used uniform cells in x_1 . The number of the cells N_x per the length $d_0 = 2\pi\sqrt{5/6} = 5.7357\cdots$ is common to all the results shown in the figures and is 2000 (note that $N = N_x$ for $d = d_0$, $N = 3N_x$ for $d = 3d_0$, etc.). Therefore, the size of a cell Δx when the oscillating plate is located at $x_1 = 0$ is $\Delta x = d_0/2000 = 0.0028\cdots$. The time step is also uniform, and the number of steps N_t per a period is from 3.2×10^5 to 1.28×10^6 . More precisely, the time step is $\Delta t = 2\pi/3.2 \times 10^5 = 1.9 \cdots \times 10^{-5}$ for the result shown in Figs. 2, 3, 5, and 6, $\Delta t = 2\pi/6.4 \times 10^5 = 9.8 \cdots \times 10^{-6}$ for Fig. 4, and $\Delta t = 2\pi/1.28 \times 10^6 = 4.9 \cdots \times 10^{-6}$ for Fig. 7. The Δx and Δt for the results shown in Figs. 8, 9, and 11 are more or less based on the data given above.

We next give some examples of the accuracy check. Hereafter, we consider the case of $\epsilon = 0.1$, $a_w = 0.1$, and $d = d_0$. In the following, we use the four grid systems:

• (Grid 1) $\Delta x = d_0/500 = 1.1 \dots \times 10^{-2}, \quad \Delta t = 2\pi/8 \times 10^4 = 7.8 \dots \times 10^{-5}$

• (Grid 2)
$$\Delta x = d_0/1000 = 5.7 \dots \times 10^{-3}, \quad \Delta t = 2\pi/1.6 \times 10^5 = 3.9 \dots \times 10^{-5}$$

• (Grid 3)
$$\Delta x = d_0/2000 = 2.8 \dots \times 10^{-3}, \quad \Delta t = 2\pi/3.2 \times 10^5 = 1.9 \dots \times 10^{-5}$$

• (Grid 4)
$$\Delta x = d_0/4000 = 1.4 \dots \times 10^{-3}, \quad \Delta t = 2\pi/6.4 \times 10^5 = 9.8 \dots \times 10^{-6}$$

Let h stand for ρ , v_1 , T, and p, and let $h^{\text{G}i}$ denote the result of h based on the grid (Grid i) (i = 1, 2, 3, 4). At $t/2\pi = 200$, we have $|h^{\text{G}1} - h^{\text{G}4}| < 1.2 \times 10^{-4}$, $|h^{\text{G}2} - h^{\text{G}4}| < 3.3 \times 10^{-5}$, and $|h^{\text{G}3} - h^{\text{G}4}| < 8.6 \times 10^{-6}$. This shows that the grid systems used in the computation have given the results of sufficient accuracy. Similar checks have been carried out by changing Δx for a fixed Δt ($\Delta t = 9.8 \cdots \times 10^{-6}$) and by changing Δt for a fixed Δx ($\Delta x = 1.1 \cdots \times 10^{-2}$). But the results are omitted here.

The total mass of the gas between the two plates per unit are of the plates is expressed as

$$m(t) = \int_{x_w(t)}^d \rho(t, x_1) dx_1,$$
 (C1)

which is constant theoretically. However, because of numerical errors, it changes slightly with time. If we plot |m(t) - m(0)|/m(0) versus t, it increases almost linearly in t, oscillating with an amplitude that does not increase with t, after some tens of the periods. We evaluated the average amplitude r_m [i.e., the average of the maximum minus the minimum in each period

 $(n < t/2\pi \le n + 1; n = 100, 101, ..., 199)$ over $100 \le t/2\pi \le 200$] and the average rate of increase s_m of the maximum per one period (i.e., the maximum in $199 < t/2\pi \le 200$ minus the maximum in $100 < t/2\pi \le 101$ divided by 100). The result is as follows: $r_m \approx 3.2 \times 10^{-6}$, $s_m \approx 3.7 \times 10^{-8}$ for (Grid 1); $r_m \approx 8.3 \times 10^{-7}$, $s_m \approx 1.8 \times 10^{-8}$ for (Grid 2); $r_m \approx 2.7 \times 10^{-7}$, $s_m \approx 9.0 \times 10^{-9}$ for (Grid 3); $r_m \approx 1.6 \times 10^{-7}$, $s_m \approx 4.5 \times 10^{-9}$ for (Grid 4). The results show that better mass conservation is attained with finer grid systems.

Finally, we consider the momentum and energy transfer. In the time-periodic state, $\overline{\mathcal{P}^{L}} = \overline{\mathcal{P}^{R}}$ and $\overline{\mathcal{E}^{L}} = \overline{\mathcal{E}^{R}}$ hold (cf. Sec. VIIB). We check how accurately these relations are satisfied numerically. Let us put $\mathcal{P}_{err} = (\overline{\mathcal{P}^{L}} - \overline{\mathcal{P}^{R}})/a_{w}$ and $\mathcal{E}_{err} = (\overline{\mathcal{E}^{L}} - \overline{\mathcal{E}^{R}})/a_{w}$. Then, $\mathcal{P}_{err} \approx 3.2 \times 10^{-6}$, $\mathcal{E}_{err} \approx 2.5 \times 10^{-4}$ (Grid 1); $\mathcal{P}_{err} \approx -3.5 \times 10^{-6}$, $\mathcal{E}_{err} \approx 2.5 \times 10^{-4}$ (Grid 2); $\mathcal{P}_{err} \approx -6.8 \times 10^{-6}$, $\mathcal{E}_{err} \approx 2.5 \times 10^{-4}$ (Grid 3); $\mathcal{P}_{err} \approx -8.2 \times 10^{-6}$, $\mathcal{E}_{err} \approx 2.5 \times 10^{-4}$ (Grid 4). This shows that the grid refinement (Grid 1) \rightarrow (Grid 4) does not improve the conservation properties of momentum and energy. On the other hand, we obtained a better result: $\mathcal{P}_{err} \approx 4.0 \times 10^{-7}$, $\mathcal{E}_{err} \approx 3.2 \times 10^{-5}$ for the grid system with $\Delta x = d_{0}/500 = 1.1 \cdots \times 10^{-2}$ and $\Delta t = 2\pi/6.4 \times 10^{5} = 9.8 \cdots \times 10^{-6}$. This and other tests suggest that to have better momentum and energy conservation, smaller time steps are required.

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